

# Measurement of $\pi^0$ photoproduction on the proton from 550 to 1500 MeV at GRAAL

O. Bartalini<sup>2,9</sup>, V. Bellini<sup>1,5</sup>, J.P. Bocquet<sup>12</sup>, M. Capogni<sup>2,9,a</sup>, L. Casano<sup>9</sup>, M. Castoldi<sup>7</sup>, P. Calvat<sup>12</sup>, A. D'Angelo<sup>2,9</sup>, R. Di Salvo<sup>9</sup>, A. Fantini<sup>2,9</sup>, C. Gaulard<sup>4,b</sup>, G. Gervino<sup>3,10</sup>, F. Ghio<sup>8,11</sup>, B. Girolami<sup>8,11</sup>, A. Giusa<sup>1,6</sup>, V. Kouznetsov<sup>13</sup>, A. Lapik<sup>13</sup>, P. Levi Sandri<sup>4</sup>, A. Lleres<sup>12,c</sup>, D. Moricciani<sup>9</sup>, A.N. Mushkarenkov<sup>13</sup>, V. Nedorezov<sup>13</sup>, L. Nicoletti<sup>2,9,12</sup>, C. Perrin<sup>12</sup>, D. Rebreyend<sup>12</sup>, F. Renard<sup>12</sup>, N. Rudnev<sup>13</sup>, T. Russew<sup>12</sup>, G. Russo<sup>1,6</sup>, C. Schaerf<sup>2,9</sup>, M.-L. Sperduto<sup>1,6</sup>, M.-C. Sutura<sup>6</sup>, and A. Turlone<sup>13</sup>

<sup>1</sup> Dipartimento di Fisica ed Astronomia, Università di Catania, via Santa Sofia 64, I-95123 Catania, Italy

<sup>2</sup> Dipartimento di Fisica, Università di Roma "Tor Vergata", via della Ricerca Scientifica 1, I-00133 Roma, Italy

<sup>3</sup> Dipartimento di Fisica Sperimentale, Università di Torino, via P. Giuria, I-00125 Torino, Italy

<sup>4</sup> INFN - Laboratori Nazionali di Frascati, via E. Fermi 40, I-00044 Frascati (Roma), Italy

<sup>5</sup> INFN - Laboratori Nazionali del Sud, via Santa Sofia 44, I-95123 Catania, Italy

<sup>6</sup> INFN - Sezione di Catania, via Santa Sofia 64, I-95123 Catania, Italy

<sup>7</sup> INFN - Sezione di Genova, via Dodecanneso 33, I-16146 Genova, Italy

<sup>8</sup> INFN - Sezione di Roma I, piazzale Aldo Moro 2, I-00185 Roma, Italy

<sup>9</sup> INFN - Sezione di Roma II, via della Ricerca Scientifica 1, I-00133 Roma, Italy

<sup>10</sup> INFN - Sezione di Torino, I-10125 Torino, Italy

<sup>11</sup> Istituto Superiore di Sanità, viale Regina Elena 299, I-00161 Roma, Italy

<sup>12</sup> IN2P3, Laboratoire de Physique Subatomique et de Cosmologie, F-38026 Grenoble, France

<sup>13</sup> Institute for Nuclear Research, 117312 Moscow, Russia

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**Abstract.** Neutral pion photoproduction has been measured from 550 to 1500 MeV with the GRAAL facility, located at the ESRF in Grenoble. Differential cross-section and beam asymmetry have been measured over a wide angular range. These high-precision data improve significantly the database for the beam asymmetry in the second and third resonance regions covering for the first time the energy domain 1100–1500 MeV. New partial-wave analyses including these data are presented for which the beam asymmetry brings crucial constraints.

**PACS.** 13.60.Le Meson production – 13.88.+e Polarization in interactions and scattering – 25.20.Lj Photoproduction reactions

## 1 Introduction

Thanks to the recent development of new accelerators worldwide, meson photoproduction (or more generally electroproduction) is nowadays a major tool to investigate  $N^*$  properties [1]. Due to its simplicity and high production rate, single-pion photoproduction plays a special role and has already been abundantly studied. Most of the  $N^*$  electromagnetic couplings have been extracted through multipole analyses of these channels [2].

<sup>a</sup> Present address: ENEA - C.R. Casaccia, via Anguillarese 301, I-00060 Roma, Italy.

<sup>b</sup> Present address: IN2P3, Centre de Spectrométrie Nucléaire et de Spectrométrie de Masse, F-91405 Orsay, France.

<sup>c</sup> e-mail: lleres@lpsc.in2p3.fr

Until recently, the database was containing essentially cross-section results often with poor statistics and limited angular range. With the onset of new facilities (Bonn-ELSA, GRAAL, JLab-CLAS, LEGS, LEPS and Mainz-MAMI), using systematically polarized beam and/or target combined with large-acceptance detectors, the experimental context has now significantly improved. In particular, the beam asymmetry  $\Sigma$ , accessible with a linearly polarized beam, has proven recently to be particularly sensitive to the resonant contributions. For instance, in the region of the  $\Delta$ -resonance, a major effort has been undertaken over the last decade to extract the tiny  $E2/M1$  ratio in the  $N$ - $\Delta$  transition for which high-precision data have been obtained [3,4]. Thanks to the beam asymmetry, an impressive accuracy has now been achieved on this ratio,

despite some model dependence in the extraction, and its value differs largely from older results. On the other hand, we have recently measured  $\Sigma$  over a large angular range for the reactions  $\gamma p \rightarrow \pi^+ n$ ,  $\gamma p \rightarrow \eta p$  and  $\gamma p \rightarrow \pi^0 \pi^0 p$  in the second and third resonance regions [5–8]. For all these channels,  $\Sigma$  provided severe constraints on the reaction mechanism. In particular,  $\Sigma$  has been the key for the quantitative determination of the  $D_{13}(1520)$  contribution in  $\eta$  photoproduction [9–12].

In the present work, we report on measurements of the beam asymmetry  $\Sigma$  and of the differential cross-section for the reaction  $\gamma p \rightarrow \pi^0 p$  over large energy ( $E_\gamma = 550\text{--}1500$  MeV,  $W = 1400\text{--}1900$  MeV) and angular ( $\theta^{\text{CM}} = 40\text{--}170^\circ$ ) ranges. For the first time, the region above 1100 MeV is extensively covered with high-precision  $\Sigma$  data.

## 2 The GRAAL set-up

### 2.1 General layout

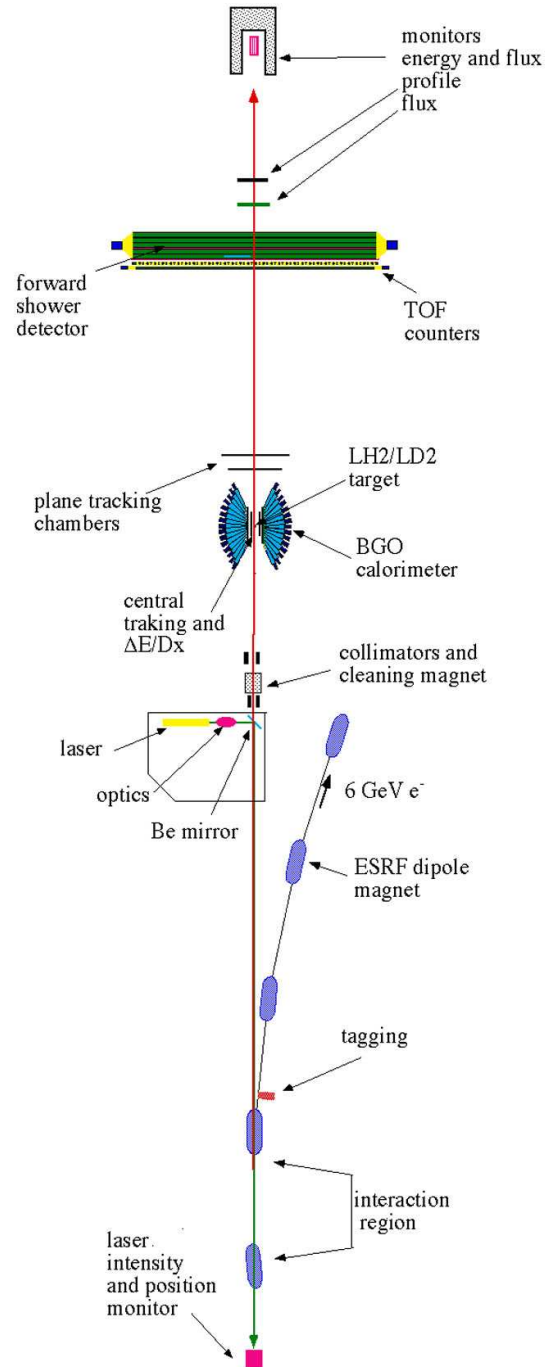
This experiment has been carried-out with the GRAAL facility (fig. 1), installed at the European Synchrotron Radiation Facility (ESRF) in Grenoble (France). The tagged and polarized  $\gamma$ -ray beam is produced by Compton scattering of laser photons off the 6.03 GeV electrons circulating in the storage ring (844 m circumference), over a 6.5 m long straight section (so-called interaction region in fig. 1). Backscattered photons of useful energies are emitted in a narrow cone and impinge on a liquid-hydrogen target placed at 41.5 m from the center of the interaction region while the scattered electrons are detected in an internal tagging system. The beam size at the target is  $\approx 2 \times 1$  cm<sup>2</sup>. A primary collimator followed by a sweeping magnet is used only to cut the low-energy tails of the beam. Two secondary collimators placed under vacuum eliminate the multitude of low-energy photons produced by showers in the main collimator and are essential to get a reduced electromagnetic background in the detector. Before reaching a SPACAL-type calorimeter (scintillating fibers embedded in a lead matrix) [13] which serves also as beam dump, photons pass through a thin monitor used to measure the beam flux.

### 2.2 Compton beam

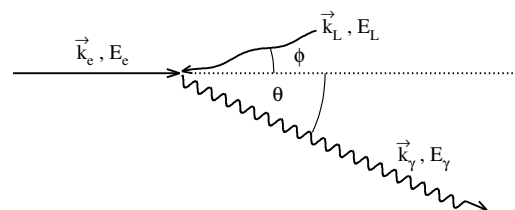
In the case of Compton scattering of low-energy photons off ultra-relativistic electrons, the energy dependence of the scattered photon *versus* angle in the laboratory frame reduces to

$$E_\gamma = \frac{4\gamma^2 E_L}{1 + \frac{4\gamma E_L}{m_e} + \theta^2 \gamma^2}, \quad (1)$$

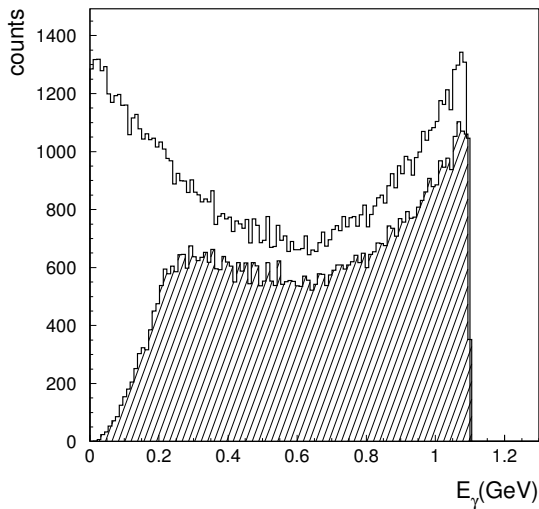
where  $\theta$  is the angle between the scattered photon and the incident direction of the electron,  $\gamma$  is the Lorentz factor of the electron ( $\approx 12000$  for the 6 GeV ESRF beam),  $E_L$  the energy of the incident photon and  $m_e$  the mass of the



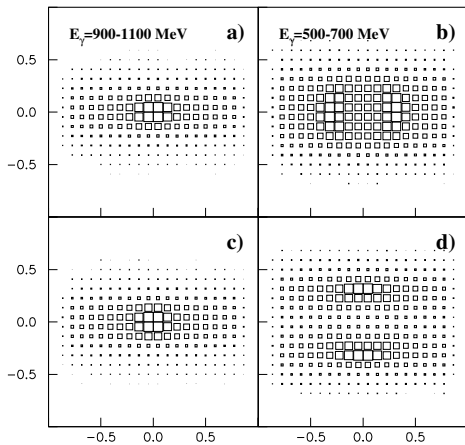
**Fig. 1.** General layout of the GRAAL set-up (the drawing is out of scale).



**Fig. 2.** Definition of Compton-scattering kinematical variables.



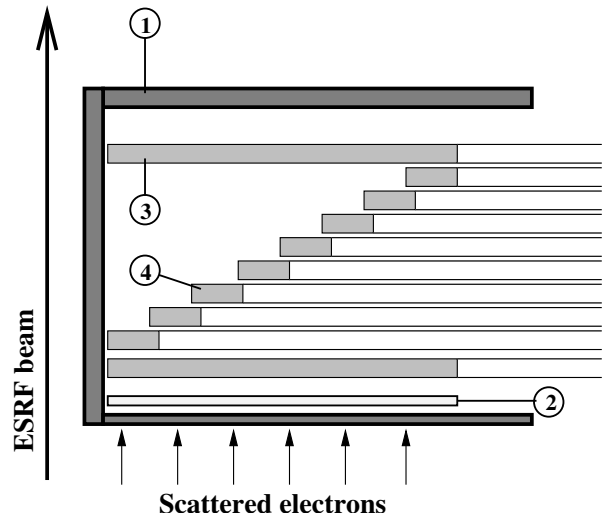
**Fig. 3.** Simulated Compton energy spectrum for  $\lambda = 514$  nm. The initial spectrum is compared to the spectrum at the target (hatched histogram). The low-energy cut is due to collimation while the difference at medium and high energy is due to the various elements on the path of the  $\gamma$ -ray beam (windows, Be mirror, ...).



**Fig. 4.** Simulated transverse beam profile at the target ( $\lambda = 514$  nm) for two photon energy ranges and for the two linear polarization states: vertical (a, b) and horizontal (c, d). The scales are in cm.

electron (fig. 2). The  $\phi$  angle between the incoming photon and the electron beam axis cannot exceed  $1^\circ$  given the geometry of the set-up and has no measurable influence. The maximum photon energy (Compton edge) depends on the laser wavelength. For the present experiment, we have used alternately the green line at 514 nm and a set of UV lines (MLUV) at 333, 351 and 364 nm produced by an Ar laser, giving 1.1, 1.4, 1.47 and 1.53 GeV Compton edges, respectively. By contrast to Bremsstrahlung beam, the energy spectrum is rather flat and the low-energy photons can be suppressed by collimation before they reach the target (fig. 3).

As illustrated in fig. 4, Compton scattering generates a beam shape strongly energy and polarization dependent.



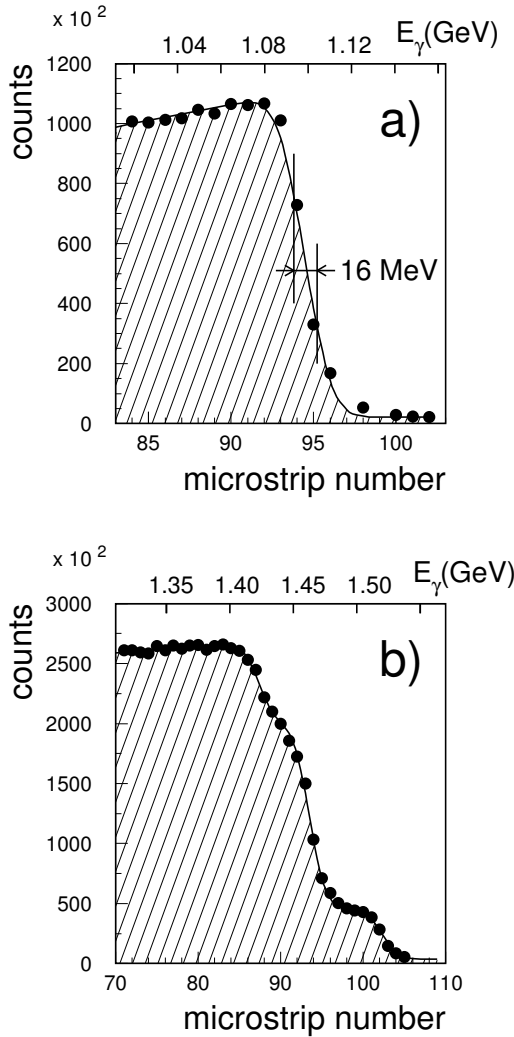
**Fig. 5.** Schematic drawing of the tagging detector: 1) X-ray shielding, 2) Si microstrip detector, 3), 4) Scintillators and light guides. See text for more details

### 2.2.1 Simulation

The  $\gamma$ -ray beam properties (resolution, polarization, intensity) depend critically on numerous optical and geometrical parameters. In order to assess their effect and optimize the beam characteristics, we have developed a complete simulation with the GEANT3 code from the CERN library. It includes a realistic description of the beam line from the interaction region to the beam dump, of the tagging system for the scattered electrons and takes precisely into account the electron beam optics and the shape of the laser beam.

### 2.2.2 Photon energy and tagging system

The photon energy is provided by an internal tagging system (fig. 5) located right after the exit dipole of the interaction region (fig. 1). The position of the scattered electron is measured by a silicon microstrip detector (128 strips with a pitch of  $300 \mu\text{m}$  and  $1000 \mu\text{m}$  thick). Two long plastic scintillators covering most of the focal plane ( $38.4$  mm) and eight additional small ones ( $6.8$  mm long) are situated behind the microstrip detector. The small scintillators overlap each of the neighbouring ones by  $2$  mm along the direction of the incoming electron beam. The signal rise time of the plastic scintillators is less than  $1.8$  ns and a custom made electronics, based on GaAs technology, provides the synchronization of the discriminated signal of one of the long scintillators with the RF signal of the electron beam. The final logical signal has a total time jitter less than  $100$  ps with respect to the timing of the electron beam and is capable of separating two electrons of adjacent bunches ( $2.8$  ns). This signal drives a triple coincidence (called “TAG” hereafter) with the second long scintillator and any of the small plastic scintillators that provides an excellent global start signal for all Time-of-Flight



**Fig. 6.** Compton edges measured by the tagging system for two different laser settings: a) green line (514 nm) and b) Multi-Line UV (364, 351, 333 nm). The three lines are clearly resolved. The solid lines represent the fitted curves from which the energy resolution (16 MeV FWHM) and the Compton-edge position (with an accuracy of  $0.03 \mu\text{strip}$ ) are extracted. See text for more details.

(ToF) measurements and enters in the different triggers (see sect. 2.5) of the experiment.

The whole system is shielded against the huge X-ray background by a  $4\pi$  box made of stainless steel and tungsten and having a total thickness equivalent to 8 mm of lead. The thickness of this box and the minimum distance allowed between its wall and the electron beam determines our threshold of 550 MeV. The high-energy part of the Compton spectrum measured by the microstrip detector is displayed in fig. 6 for the green and MLUV settings of the laser. As illustrated in this figure, the experimental energy resolution can be directly extracted from the fit of the Compton edge with an error function. The measured resolution of 16 MeV (FWHM) is in agreement with the estimated one from the simulation and is dominated

by the energy spread of the electrons in the storage ring (FWHM = 14 MeV). In the case of MLUV, the three edges are clearly resolved and their peak position and relative intensity can be extracted. The peak position of a given line is measured with a high accuracy ( $9 \mu\text{m}$  equivalent to 0.3 MeV).

The relation between the  $\gamma$ -ray energy  $E_\gamma$  and the distance,  $x_e$ , of the scattered electron to the beam axis can be approximated by the following expression:

$$E_\gamma = E_e \frac{x_e}{a_0 + x_e}, \quad (2)$$

where  $E_e$  is the electron beam energy and  $a_0$  is a geometrical factor, constant to an excellent approximation over the tagged energy range ( $da_0/a_0 = 5 \cdot 10^{-4}$ ).

The beam energy given by the ESRF machine group, based on a  $\int Bdl$  measurement over the ring, was:

$$E_e^{\text{ESRF}} = 6040 \pm 20 \text{ MeV}.$$

We have performed two independent measurements of this energy. First, by using the precisely known energy threshold of the reaction  $\gamma p \rightarrow \eta p$ , we have obtained a consistent value with an improved accuracy [14]:

$$E_e^\eta = 6030.6 \pm 6.0 \text{ MeV}.$$

More recently, taking advantage of our ability to measure the transverse polarization of the stored electrons (Sokolov-Ternov effect), using the beam profiler of the GRAAL set-up, we have achieved an excellent precision on the beam energy by means of the resonant depolarization method [15]:

$$E_e^{\text{RD}} = 6027.6 \pm 1.3 \text{ MeV}.$$

The geometrical factor  $a_0$  depends on the radius of curvature in the dipole and on the longitudinal position of the microstrip detector. Its value can be inferred from the distance between Compton edges associated with different laser lines. We have obtained:

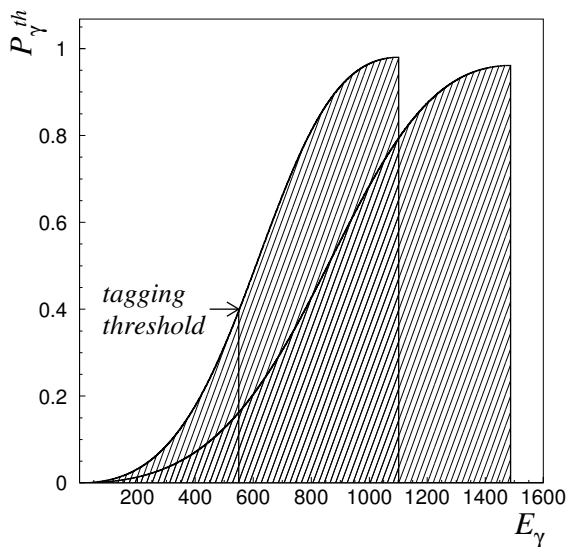
$$a_0 = 159.92 \pm 0.3 \text{ mm},$$

Finally, the distance  $x_e$  of the scattered electron to the beam axis is

$$x_e = (n_{\mu\text{strip}} - 0.5)\Delta x + x_0, \quad (3)$$

where  $n_{\mu\text{strip}}$  is the number of the hit microstrip,  $\Delta x$  the pitch ( $300 \mu\text{m}$ ) and  $x_0$  the distance of the first microstrip to the stored electron beam. This latter parameter is deduced from the fit of the position of the Compton edge run by run.

Taking into account all possible sources of uncertainties, we estimate that the absolute error on the photon energy calibration does not exceed  $\approx 2 \text{ MeV}$  over the whole tagged range.



**Fig. 7.** Theoretical linear polarization of the  $\gamma$ -ray beam for two laser lines:  $\lambda = 514$  nm (green line) and  $\lambda = 351$  nm (UV line). The energy threshold (550 MeV) introduced by the tagging system is shown.

### 2.2.3 Beam polarization

The  $\gamma$ -ray beam polarization is calculated with the help of the Klein-Nishina formula and was first derived for such asymmetric collisions by Arutyunian *et al.* [16]. Its energy dependence is displayed in fig. 7 for two of the used wavelengths (514 nm, 351 nm) and in the case of a 100% linearly polarized laser beam. The  $\gamma$ -ray polarization is close to 1 at the Compton edge (98% for green and 96% for UV) and decreases smoothly with energy down to a minimum of  $\approx 20\%$  at our energy threshold ( $E_\gamma \approx 550$  MeV).

Below 1100 MeV, an asymmetry measurement can be performed with any of the two laser lines. This possibility has been used to perform an experimental cross-check of the polarization energy dependence presented in sect. 3.5.3.

Polarization of backscattered photons is proportional to the incident laser light polarization. Before each period of data taking, the laser polarization was measured into the laser hutch, after all optical elements. In order to avoid any distortion during the beam transport, a special viewport was used to minimize mechanical stress on the window. More recently, a new polarimeter has been installed after the intersection region into the ESRF tunnel and has allowed to confirm the previous measurements. We used a conservative value for the laser polarization and its associated error:

$$P_L = 98 \pm 2\%.$$

Polarization calculation at one given energy—or equivalently at one given scattering angle  $\theta$ —results from an integration over the azimuthal angle of all photons emitted at the same  $\theta$ :

$$P(\theta) = \frac{\int P(\theta, \varphi) \frac{d\sigma}{d\Omega}(\theta, \varphi) d\varphi}{\int \frac{d\sigma}{d\Omega}(\theta, \varphi) d\varphi}. \quad (4)$$

With the help of the simulation, we have determined that a broad collimation ( $15 \times 12$  mm<sup>2</sup>) was compatible with the target diameter (4 cm) while preserving the polarization of the low-energy tagged photons, even in the case of small drifts of the beam position.

Due to the scattering of electrons on the residual vacuum over the 6.5 m long interaction region, unpolarized Bremsstrahlung photons contaminate the Compton beam and affect the average degree of polarization. During the experiment, short periods of data acquisition were taken without laser (typically 5 min every 50 min) in order to measure the Bremsstrahlung background alone. Thanks to an excellent vacuum in the ring, the fraction of Bremsstrahlung photons  $F_{\text{Brems}}/F_{\text{Compton}}$  remained of the order of 0.1% over the entire experiment.

In summary, we have used the following expression to calculate the beam polarization and its associated error:

$$P_\gamma = P_L P_\gamma^{\text{th}} \left( 1 - \frac{F_{\text{Brems}}}{F_{\text{Compton}} + F_{\text{Brems}}} \right), \quad (5)$$

$$\frac{\delta P_\gamma}{P_\gamma} = \sqrt{\left( \frac{\delta P_L}{P_L} \right)^2 + \left( \frac{\delta P_\gamma^{\text{th}}}{P_\gamma^{\text{th}}} \right)^2 + \left( \frac{\delta F_{\text{Brems}}}{F_{\text{Compton}} + F_{\text{Brems}}} \right)^2}, \quad (6)$$

where  $P_\gamma^{\text{th}}$  is the calculated polarization. In the MLUV case, we took into account the superposition of the three laser lines with the measured relative intensities as described previously.

In practice, the only significant source of error came from the laser beam polarization:

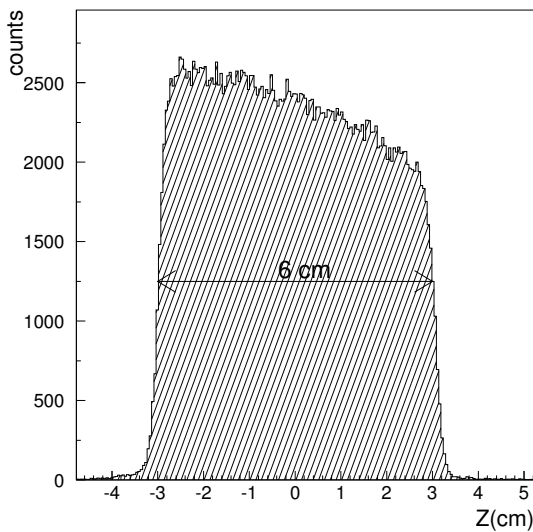
$$\frac{\delta P_\gamma}{P_\gamma} = \frac{\delta P_L}{P_L} = 0.02. \quad (7)$$

### 2.2.4 Flux monitoring

To avoid pile-up problems at high counting rate, the  $\gamma$ -ray beam intensity has been measured by means of a thin monitor. It consists of three plastic scintillators ( $5 \times 5 \times 0.5$  cm<sup>3</sup>), the first one acting as a veto, the two others being in coincidence, and a thin Al converter (2 mm) placed after the veto. Its efficiency ( $2.68\% \pm 0.03$ ) has been measured by comparison with the SPACAL absorber at low rate. It takes into account the  $\gamma$ -ray conversion between the target and the thin monitor and it is found experimentally independent of the  $\gamma$ -ray energy in the covered range. The typical flux during the experiment was  $10^6$  Hz.

## 2.3 The LH2 target

The target cell is an aluminum hollow cylinder of 4 cm in diameter closed by thin mylar windows (100  $\mu$ m) at both ends. To satisfy different experimental requirements (resolution *versus* counting rates), three different lengths (3,



**Fig. 8.** Longitudinal target profile reconstructed from the vertex distribution of the  $p\pi^+\pi^-$  reaction in the case of a 6 cm long target.

6 and 12 cm) have been used over the data taking periods. When filled with hydrogen, the working temperature is 18 K, giving a density  $\rho \approx 7 \cdot 10^{-2} \text{ g/cm}^3$ .

The longitudinal position of the target is a crucial parameter for the control of the acceptance. The target profile displayed in fig. 8 corresponds to the vertex distribution (along the beam axis) of the three charged particles reaction  $p\pi^+\pi^-$ , provided by the central tracking detector (see sect. 2.4). From this distribution, the exact target position could be extracted. It should be noted that the profile shape reflects the geometrical acceptance of the central part of the detector for the selected events.

## 2.4 The LA $\gamma$ RANGE detector

The 4 $\pi$  LA $\gamma$ RANGE detector (fig. 1) has been designed to detect both neutral and charged particles produced in photoproduction reactions on light nuclei with total center-of-mass energy  $W \leq 2 \text{ GeV}$ . The apparatus is composed of two main parts: a central one ( $25^\circ \leq \theta \leq 155^\circ$ ) and a forward one ( $\theta \leq 25^\circ$ ). In the following, all resolutions are given as Full Width Half-Maximum (FWHM).

*Proton detection.* The track of protons was measured by a set of Multi-Wire Proportional Chambers (MWPC). Two cylindrical chambers with stripped cathodes were used to cover the central region and gave a reconstruction efficiency  $\geq 90\%$  with a resolution of  $3.5^\circ$  in  $\theta$  and  $4.5^\circ$  in  $\varphi$ . The forward angles were measured by two planar chambers (efficiency  $\geq 99\%$ ), each composed of two wire planes; the average polar and azimuthal resolutions were, respectively,  $1.5^\circ$  and  $2^\circ$ . Detailed discussions of the characteristics and performances of these detectors can be found in [17].

Particle identification was obtained by  $dE/dx$  technique in the central region thanks to a plastic scintillator

barrel (32 bars, 5 mm thick, 43 cm long and  $\approx 20\%$  energy resolution) situated between the cylindrical chambers and the BGO calorimeter. For the more energetic particles emitted in the forward direction, a Time-of-Flight measurement was used, provided by a double plastic scintillator hodoscope ( $300 \times 300 \times 3 \text{ cm}^3$ ) placed at a distance of 3 m from the target and having a resolution of 600 ps. This latter detector provides also a measure of the energy loss.

For the cross-sections, due to large uncertainties on the cylindrical chambers efficiency, the proton direction was deduced from the association between the scintillator barrel and the BGO calorimeter (resolution  $\sim 10^\circ$  in  $\theta$  and  $\varphi$ ).

*$\gamma$ -ray detection.* The two  $\gamma$ -rays from the  $\pi^0$  decay (branching ratio of 98.8%) were detected in the BGO calorimeter made of 480 ( $15\theta \times 32\varphi$ ) crystals, each of 21 radiation lengths so that the associated electromagnetic shower was fully contained into the calorimeter.  $\gamma$ -rays were identified as clusters of adjacent crystals (3 on average for an energy threshold of 10 MeV per crystal) with no associated hit in the barrel. The measured photon energy resolution was 3% on average and the angular resolution was, respectively,  $6^\circ$  and  $7^\circ$  for polar and azimuthal angles ( $E_\gamma \geq 200 \text{ MeV}$  and  $l_{\text{target}} = 3 \text{ cm}$ ). The response of the BGO calorimeter has been investigated in numerous publications [18–20].

At forward angles,  $\gamma$ -rays can be detected in a lead-scintillator sandwich ToF wall [21] located behind the double plastic hodoscope. This detector provides a good angular resolution but no energy measurement and for the present reaction extends only marginally the covered angular range. For the sake of simplicity, it has not been used in the present analysis.

Backward angles ( $\theta \geq 155^\circ$ ) are covered by a lead-scintillator sandwich disk not used in the present analysis.

## 2.5 Data acquisition and trigger

A specific data acquisition system [22] was used to collect the flow of events on DLT tapes at a typical rate of 200 events/s with a negligible dead time. Two main triggers were necessary for this experiment. The first one (“physical” events) was a coincidence between the “TAG” signal (see sect. 2.2.2) and an energy deposition larger than  $\sim 180 \text{ MeV}$  in the BGO calorimeter. This open trigger allowed the simultaneous detection of all meson production channels over the covered energy range and was sufficient to reject the electromagnetic background generated in the target. The second one (monitor events), corresponding to the coincidence between the thin monitor and the “TAG” signals, was recorded in parallel with a downscaling factor of 1000. This latter trigger entered in the measurement of the integrated flux.

## 3 Data analysis

The results presented hereafter correspond to the summation of several data taking periods (nine spread over

four years) with slightly different experimental conditions (laser wavelength, target length, plane chamber set-up, ...). More details can be found in [14].

### 3.1 Simulation

A complete simulation of the apparatus was developed using the GEANT3 code from the CERN library. This software includes the dependence of beam shape on energy and polarization, a realistic event generator [23] and a complete description of the different detection components of the LA $\gamma$ RANGE set-up (geometry and response). For each data taking period, a specific version adapted to the corresponding configuration was implemented.

In addition to the usual calculation of the efficiencies entering in the differential cross-section measurements, this simulation was an essential tool to determine the best set of selection cuts and their optimization *versus* background rejection and/or efficiency.

### 3.2 Tagging analysis

After checking that the correct information (TDC, ADC) for the hit plastics was available, only those events with one or two adjacent small plastics and a microstrip signal in geometrical coincidence were selected. This same selection criterion was applied to “physical” and monitor events.

### 3.3 Flux measurement

The flux measurement was based on a 32 bits scaler which counted the total number of coincidences,  $N_{\text{scal}}$ , between the thin monitor and the “TAG” signals. The integrated beam flux has been calculated for each considered energy bin  $i$  using the following formula:

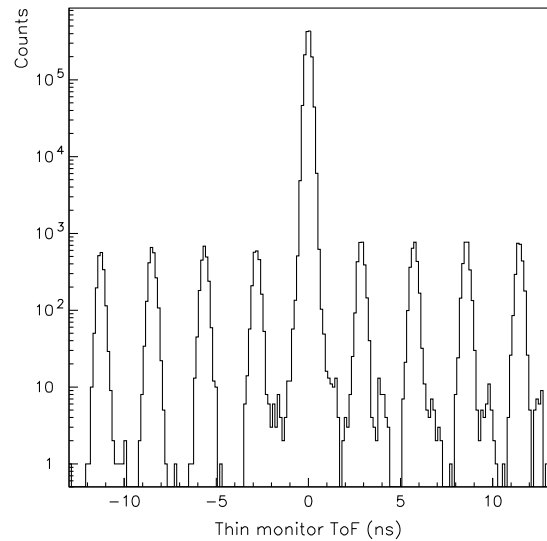
$$F(i) = N_{\text{scal}} \frac{N_{\text{mon}}^*(i)}{N_{\text{mon}}} \frac{1}{\epsilon_{\text{mon}}}, \quad (8)$$

where  $\epsilon_{\text{mon}}$  is the intrinsic detection efficiency of the thin monitor,  $N_{\text{mon}}$  and  $N_{\text{mon}}^*$  are the numbers of thin monitor triggers before and after the event analysis, *i.e.* tagging analysis and selection of the “true” peak in the ToF spectrum (see fig. 9).

A typical ToF spectrum obtained for the thin monitor is shown in fig. 9. The peaks corresponding to the different electron bunches (separated by 2.8 ns) are clearly resolved. The contribution of accidental coincidences in the main peak is negligible ( $< 0.15\%$ ).

### 3.4 Channel selection

Thanks to the  $4\pi$  geometry of the LA $\gamma$ RANGE apparatus, all final-state particles were detected. The kinematics of this two-body reaction was therefore overdetermined. We



**Fig. 9.** Typical experimental Time-of-Flight spectrum for the thin monitor. The distance between two adjacent peaks is 2.8 ns. The level of accidental coincidences in the main peak is less than 0.15%.

took advantage of this characteristic to perform channel selection by means of kinematical cuts, in contrast to the usual background subtraction method. The same method was applied for  $\pi^0 p$  and  $\eta p$  channels [24].

The first step of the analysis consisted in selecting events with only two neutral clusters in the BGO calorimeter and a single charged-particle track (see sect. 2.4). Cluster information was combined first to identify a  $\pi^0$  by means of the usual invariant mass:

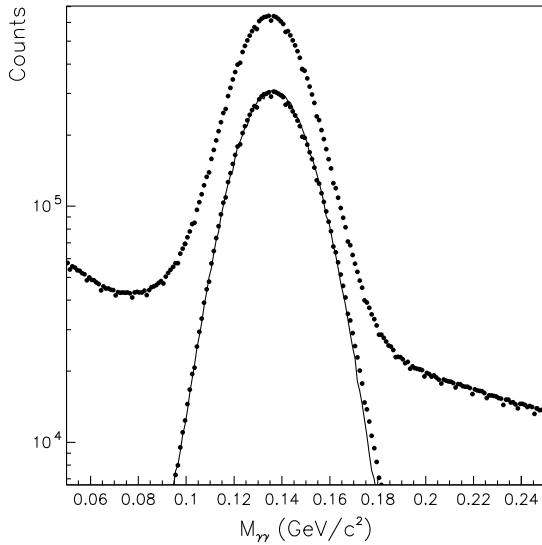
$$M_{\gamma\gamma} = \sqrt{2E_{\gamma 1}E_{\gamma 2}(1 - \cos(\theta_{\gamma 1\gamma 2}))}, \quad (9)$$

where  $\theta_{\gamma 1\gamma 2}$  is the angle between the two decay photons, and, second, to reconstruct its direction ( $\theta_{\pi^0}$ ,  $\varphi_{\pi^0}$ ) and energy  $E_{\pi^0}$ . After a cut on  $M_{\gamma\gamma}$ , we used the following set of selection cuts based on the two-body kinematics:

- $R_{\pi^0} = E_{\pi^0}/E_{\pi^0}^*$ ,
- $\Delta\theta_p = \theta_p^* - \theta_p$ ,
- $\Delta\varphi_p = \varphi_p^* - \varphi_p$ ,
- $\Delta t_p = \text{ToF}_p^* - \text{ToF}_p$  (only at forward angles),

where the “\*” quantities were calculated from the kinematics using only two pieces of experimental information: the photon energy  $E_\gamma$  and the  $\pi^0$  polar angle  $\theta_{\pi^0}$ .

The experimental distributions of  $M_{\gamma\gamma}$  are displayed in fig. 10 (closed circles). The upper spectrum shows the invariant mass after the first step of the analysis and the lower one after applying all selection cuts. One can see that the initial condition (two neutral clusters and one track) is already quite selective thanks to the  $4\pi$  coverage of the apparatus. The latter spectrum is compared with the simulation of the  $\pi^0 p$  reaction. The excellent agreement underlines the reliability of the simulation code and the low level of background (hadronic and electromagnetic).



**Fig. 10.**  $\pi^0$  invariant-mass spectrum. Data (closed circles) before and after all kinematical cuts are compared to simulation (solid line). See text for more details.

The small contribution of the hadronic background has been estimated with the simulation. Mainly composed of two-pions production ( $\pi^+\pi^0n$  and  $\pi^0\pi^0p$ ), it has been found below 1% over the whole tagged range. Contributions of the target cell have been measured with empty target runs and amount to 1.3% in accordance with the walls thickness. Finally, accidental coincidences have been found to be negligible ( $< 0.1\%$ ).

### 3.5 Measurement of $\Sigma$

#### 3.5.1 Method

For a linearly polarized beam with polarization  $P_\gamma$ , the differential cross-section at a given photon energy can be written as

$$\frac{d\sigma}{d\Omega}(\theta, \varphi) = \left( \frac{d\sigma}{d\Omega}(\theta) \right)_0 [1 + P_\gamma \Sigma(\theta) \cos 2(\varphi - \varphi_\gamma)], \quad (10)$$

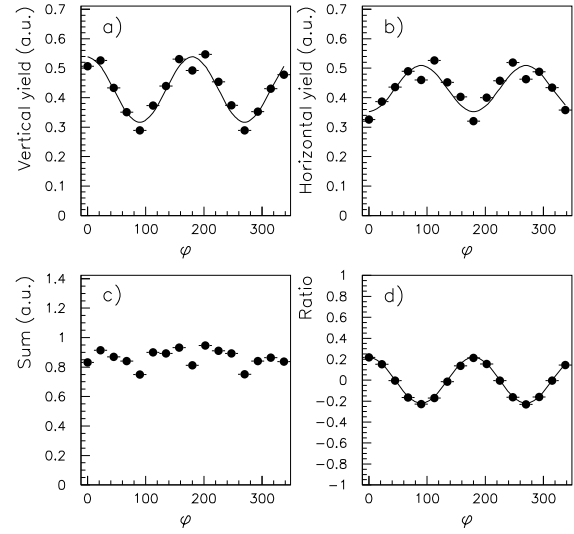
where  $(d\sigma/d\Omega)_0$  is the unpolarized differential cross-section,  $\varphi$  the azimuthal angle of the reaction plane and  $\varphi_\gamma$  the direction of the photon electric field. Alternating periods with vertical and horizontal polarization, we have obtained, for each energy and polar angular bin, the azimuthal yields normalized by the integrated flux,  $\tilde{N}_V$  and  $\tilde{N}_H$ :

$$\tilde{N}_V(\varphi) = \frac{N(\varphi, \varphi_\gamma = 90^\circ)}{F_V} \quad (11a)$$

$$= \tilde{N}_{\epsilon_V}(\varphi)[1 + P_\gamma \Sigma \cos(2\varphi)], \quad (11b)$$

$$\tilde{N}_H(\varphi) = \frac{N(\varphi, \varphi_\gamma = 0^\circ)}{F_H} \quad (12a)$$

$$= \tilde{N}_{\epsilon_H}(\varphi)[1 - P_\gamma \Sigma \cos(2\varphi)], \quad (12b)$$



**Fig. 11.** Azimuthal normalized yields: a) vertical polarization; b) horizontal polarization; c) sum of a) and b) (eq. 13); d) ratio a)-b)/c) (eq. 14). The solid lines represent the fitted curves.

where  $F_{V,H}$  and  $\epsilon_{V,H}$  are the integrated flux and azimuthal efficiencies for the two polarization states, respectively. By combining eqs. (11b) and (12b) and assuming that  $\epsilon(\varphi)$  does not depend on the state of polarization, we can then calculate the sum and the usual ratio:

$$\tilde{N}_V(\varphi) + \tilde{N}_H(\varphi) = \tilde{N}\epsilon(\varphi), \quad (13)$$

$$\frac{\tilde{N}_V(\varphi) - \tilde{N}_H(\varphi)}{\tilde{N}_V(\varphi) + \tilde{N}_H(\varphi)} = P_\gamma \Sigma \cos(2\varphi). \quad (14)$$

The experimental distributions corresponding to eqs. (11) to (14), summed over all energies and polar angles, are displayed in fig. 11. Thanks to the cylindrical symmetry of our set-up, the complete azimuthal distribution of events is observed and  $\tilde{N}_V$  as well as  $\tilde{N}_H$  exhibit the correct  $\varphi$  dependence (figs. 11a, b). The azimuthal acceptance is given by the sum spectrum (fig. 11c) which corresponds to the response of the apparatus to an unpolarized beam (same number of photons in the two orthogonal states); the observed anisotropies (of the order of 10%) are mostly due to the mechanical structure of the BGO calorimeter. As expected, the ratio (fig. 11d), from which the efficiency drops out, allows to get rid of most of the systematic errors as evidenced by an improved azimuthal distribution. Accordingly, we have fitted the ratio by the function  $a \cos(2\varphi)$  (eq. 14) to extract  $P_\gamma \Sigma$  and then  $\Sigma$  using the known energy dependence of  $P_\gamma$ . All reduced  $\chi^2$  values were found close to 1.

As mentioned earlier, we have implicitly assumed that the azimuthal efficiency  $\epsilon(\varphi)$  is the same for both polarization states. This is not necessarily true since the beam shape is different for horizontal and vertical polarizations (see fig. 4). With the help of the detector simulation, we have investigated the effect of the vertex position on the azimuthal efficiency. We have found an efficiency variation



less than 1% whatever the transverse position and concluded that this is a negligible effect.

### 3.5.2 Corrections and errors

Among the various factors that can alter the measured asymmetries, only two were found non-negligible: the  $\varphi$  binning and the hadronic background. The first one tends to lower the true asymmetry and can be readily calculated whereas the latter one depends on the nature (asymmetry) of the background and has to be estimated by the simulation. Given the statistics and the  $\varphi$  resolution, the optimal  $\varphi$  binning was 16; the corresponding reduction factor  $R_\varphi$  was 2.6%. For the background contribution, we made the assumption of a zero asymmetry whatever the bin. In such a case, the correction factor is equal to the background/signal ratio  $R_b$ , which never exceeded 2%. The two corrections were combined as follows:

$$\Sigma_{\text{true}} = \Sigma_{\text{meas}}(1 + R_\varphi)(1 + R_b). \quad (15)$$

Two sources of systematical errors were considered: the uncertainty on the laser beam polarisation ( $\delta\Sigma/\Sigma = 2.0\%$ ) and on the hadronic-background contributions ( $\delta\Sigma \leq 0.02$ ). For this second one, we took into account an uncertainty on the amount of background  $R_b$  and also on the assumed value of the background asymmetry. The absolute statistical uncertainty on  $\Sigma$ , given by the fitting procedure, was found in most cases lower than 0.03. All systematical and statistical errors have been summed quadratically.

### 3.5.3 Period comparison

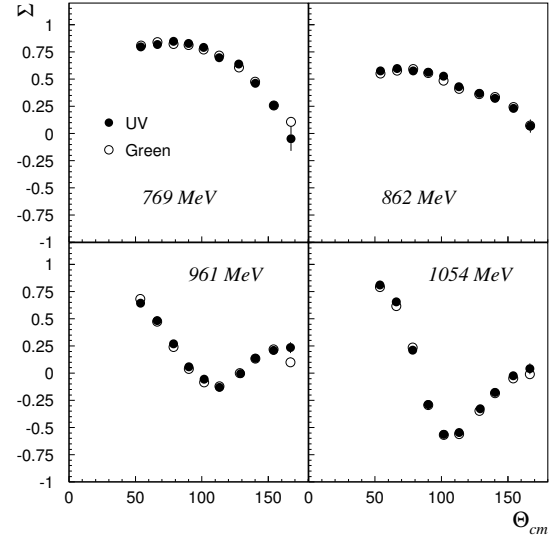
Special care was taken to check the quality and stability of the analysis before merging data from the nine different periods. This was done first by monitoring the response of each detector of the GRAAL apparatus over the whole sample of selected events. The measured quantities ( $\Sigma$  and  $d\sigma/d\Omega$ ) were also used to cross-check the various periods by calculating a  $\chi^2$  between each period and the global result. Despite the small error bars, the obtained reduced  $\chi^2$  were all found close to 1 and the final values of the beam asymmetry were extracted from the whole set of data.

Comparing results from green and UV periods allowed also to perform a crucial test on the energy dependence of the  $\gamma$ -ray beam polarization. As seen in fig. 7, below 1100 MeV,  $\Sigma$  could be measured with either of the two lines. The comparison, displayed in fig. 12, shows an excellent agreement and confirms that the beam polarization was well under control.

## 3.6 Measurement of $d\sigma/d\Omega$

### 3.6.1 Method

The differential cross-section for a given bin in energy  $E_\gamma$  and in  $\cos\theta_{\text{cm}}$ , with  $\theta_{\text{cm}}$  the  $\pi^0$  center-of-mass angle, was



**Fig. 12.** Comparison of the beam asymmetry obtained for two laser lines for four energies in the overlapping region.

calculated using the following expression:

$$\frac{d\sigma}{d\Omega}(\cos\theta_{\text{cm}}, E_\gamma) = \frac{N(\cos\theta_{\text{cm}}, E_\gamma)}{\epsilon(\cos\theta_{\text{cm}}, E_\gamma)F(E_\gamma)\rho l \Delta\Omega}, \quad (16)$$

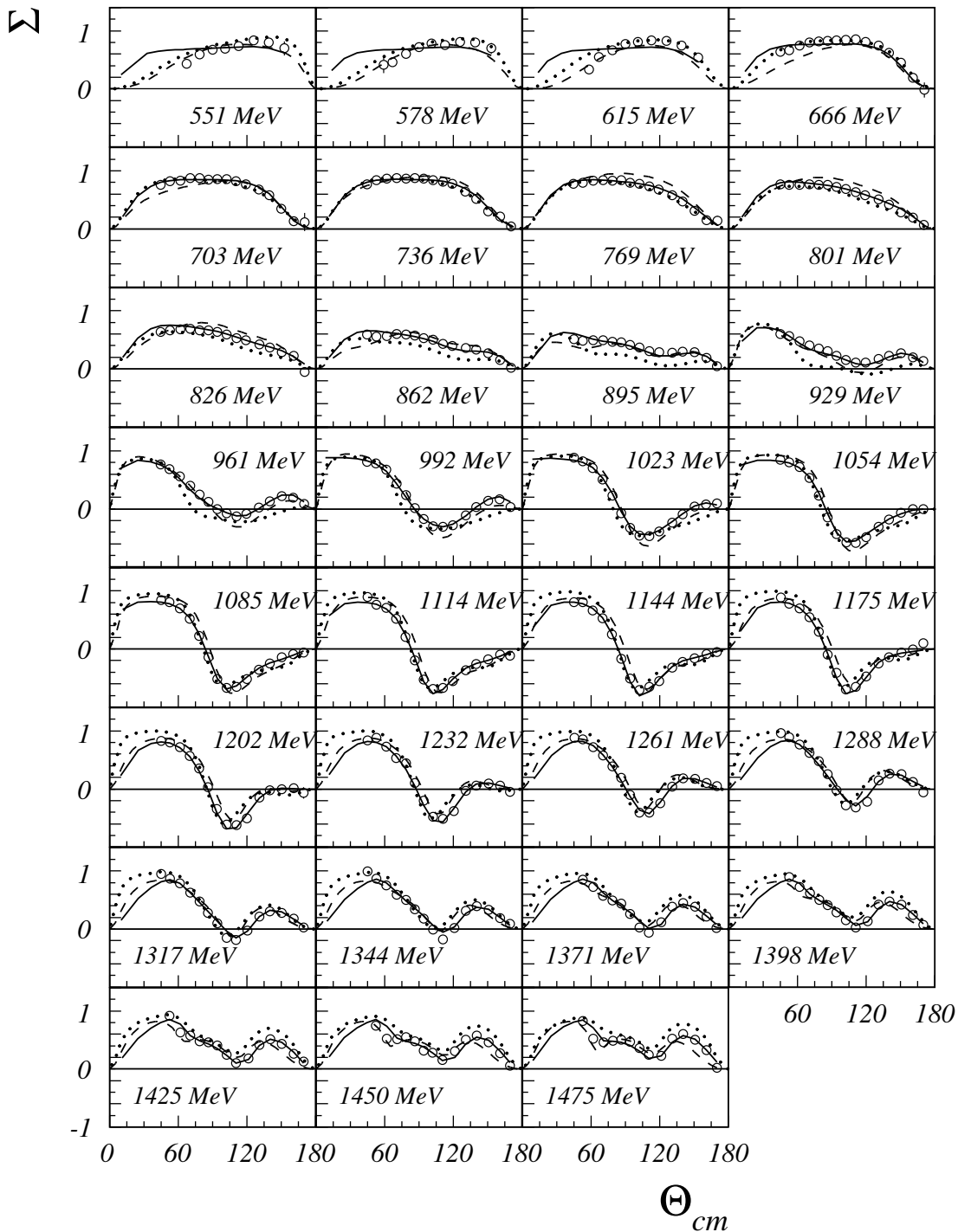
where  $N$  is the number of selected events,  $\epsilon$  the detection efficiency,  $F$  the integrated beam flux,  $\rho$  the hydrogen density,  $l$  the target length and  $\Delta\Omega$  the solid angle corresponding to the chosen  $\cos\theta_{\text{cm}}$  binning (20 bins in the present case —  $\Delta\Omega = 0.2\pi$ ). The values of  $\rho$ ,  $l$  and  $F$  have been discussed previously (see sects. 2.3 and 3.3).

The detection efficiency  $\epsilon$  has been derived from the simulation and was the most critical quantity in this formula. Applying broad cuts at  $\pm 4\sigma$  on the five kinematical variables used for event selection, we obtained a global efficiency (including acceptance, detection, identification and selection) remaining larger than 45% over the whole energy range. To check the reliability of the simulation, a systematic comparison between experimental and simulated distributions of the different selection variables was performed. This comparison emphasized small discrepancies which have been attributed to imperfections of the simulation. Indeed, some details of the mechanical structure of the detector were not included and the associated efficiency losses have been estimated and eventually corrected.

The final cross-section values have been obtained after applying corrections for hadronic background ( $\leq 2\%$ , see sects. 3.4 and 3.5.2).

### 3.6.2 Systematical errors

Two types of systematical errors were taken into account: global normalization errors and bin-dependent errors. The global normalization error included the uncertainties on



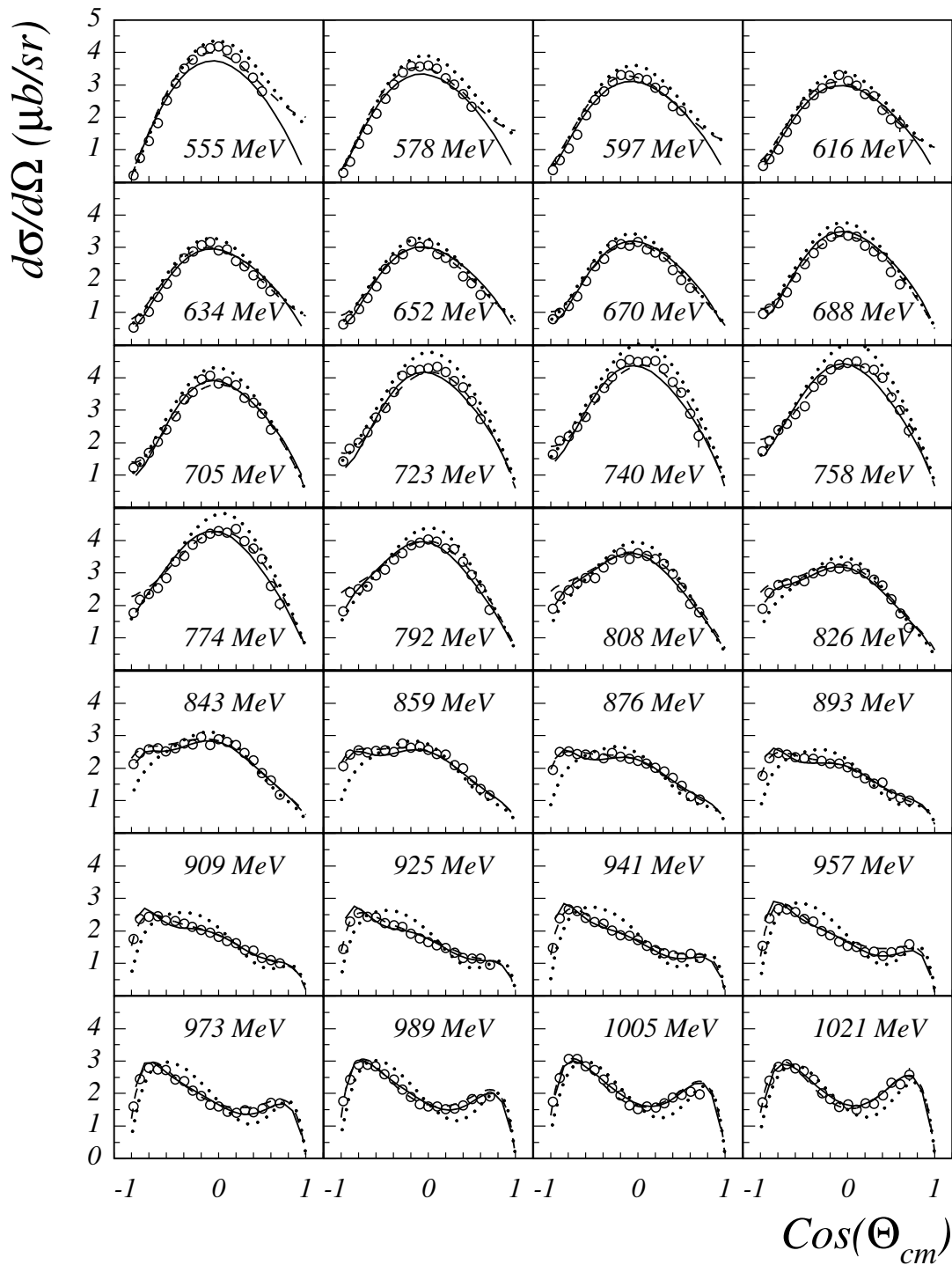
**Fig. 13.** Angular distributions of the beam asymmetry. Data are compared to the three models: SAID-FA04 (dashed line), MAID2005 (dotted line) and Bonn2005 (solid line).

the monitor efficiency, the hydrogen density and on the target length. The quadratic sum of these different contributions gave a global normalization error of 2.3%. The bin-dependent errors included uncertainties on the background contributions ( $\leq 0.8\%$ ), the efficiency ( $\leq 2\%$ ), the longitudinal target position ( $\leq 0.15 \mu\text{b}/\text{sr}$ ) and on the geometrical corrections (2–4% for a few bins). These various

bin-dependent contributions were summed quadratically with the statistical errors.

### 3.6.3 Period comparison

Differential cross-sections have been extracted from the same sample of selected runs as the beam asymmetry. A



**Fig. 14.** Differential cross-sections for energies ranging from 555 MeV to 1021 MeV. Definition of the curves as in fig. 13.

similar  $\chi^2$  procedure was also used to test the compatibility of the different periods of data taking. Because  $d\sigma/d\Omega$  is an absolute measurement, all factors that enter in the calculation must be precisely taken into account; this comparison, as well as the careful monitoring of the apparatus response, was therefore much more stringent than in the case of the asymmetry. Nevertheless, the obtained reduced  $\chi^2$  did not reveal any anomalous or bad period and the

$d\sigma/d\Omega$  values have been extracted by summing up all periods.

#### 4 Results and discussions

The complete set of asymmetry and differential cross-section data totalizing 1302 points is displayed in figs. 13

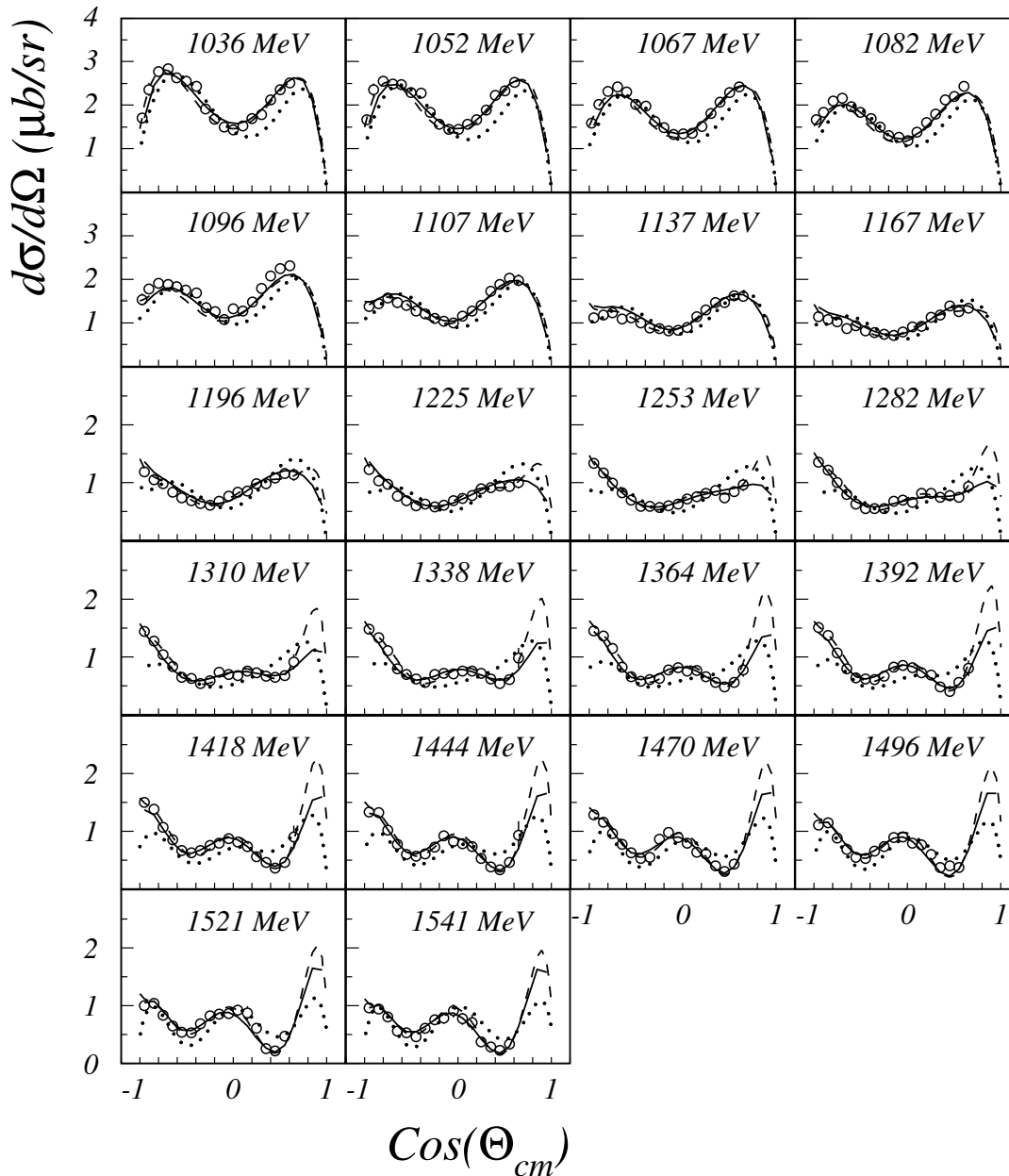


Fig. 15. Differential cross-sections for energies ranging from 1036 MeV to 1541 MeV. Curves as in fig. 14.

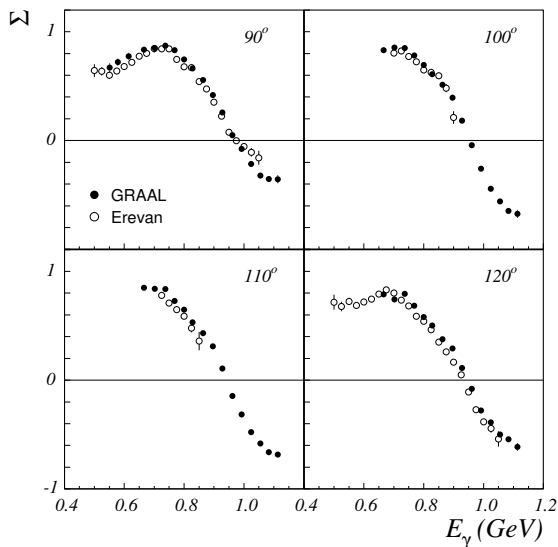
to 15. Numerical values are listed in tables 1 to 7. Error bars are the quadratic sum of statistical and systematic errors. For cross-sections, the global normalization uncertainty of 2.3% is not included. Data are available via internet [25].

#### 4.1 Comparison with previous results

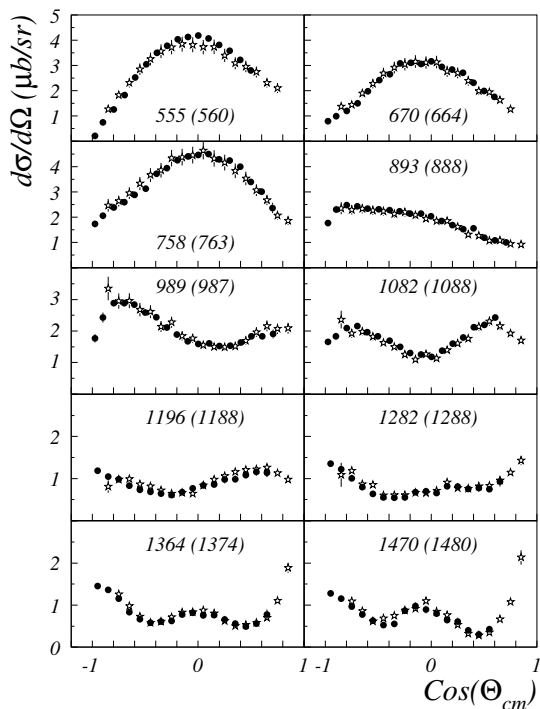
##### 4.1.1 Beam asymmetry

Apart from a few data points obtained in the seventies, the only significant results concerning the beam asymmetry have been published by the Erevan group in the energy range 500–1100 MeV and for  $\pi^0$  angles between 85 and 125° [26]. The comparison at four angles between

Erevan (open circles) and GRAAL results (closed circles) is displayed in fig. 16. At first sight, the overall agreement between the two sets of data looks satisfactory. Nevertheless, looking in more detail, a small discrepancy is observed at 110 and 120° which is attributed to a shift in energy ( $\Delta E \approx 15$  MeV) rather than to the beam polarization. Indeed, the shift remains constant over the whole energy range and in particular at the zero crossing where beam polarization is irrelevant. For the Erevan experiment, the photon energy was reconstructed from the kinematics of the proton in the magnetic spectrometer and different kinematic settings were used to cover the whole range. In our case, we used a tagged beam with one energy calibration and all angles were measured simultaneously in a  $4\pi$  detector.



**Fig. 16.** Comparison between GRAAL (closed circles) and Erevan (open circles) beam asymmetries for four angles as a function of  $E_\gamma$ .



**Fig. 17.** Comparison between GRAAL (closed circles) and Bonn (stars) differential cross-sections for the closest energy bins of the two experiments. Energies are in MeV (Bonn values are in parenthesis).

#### 4.1.2 Cross-section

Although less sparse than the beam asymmetry data, the differential cross-section database was limited until the recently published CB-ELSA results [27] which cover the energy range 0.3–3.0 GeV. The GRAAL data are absolute measurements including flux normalization in contrast to

the Bonn results where the SAID partial-wave analysis was used for absolute normalization. The comparison between the GRAAL and Bonn data is shown in fig. 17 for the closest energy bins. The overall agreement is excellent over the whole range. It should be noted that our angular range is complementary to the Bonn one, extending the range to more backward angles.

## 4.2 Discussion

We have compared our results with three models (see figs. 13-15): the SAID partial-wave analysis (FA04, dashed line) [28,29], the unitary isobar model MAID (MAID 2005, dotted line) [30–32] and the recent coupled-channel partial-wave analysis developed at Bonn (solid line) [33, 34].

The SAID partial-wave analysis for pion photoproduction uses as database the total and differential cross-sections as well as the beam and target asymmetries for the  $\pi^0 p$ ,  $\pi^+ n$ ,  $\pi^- n$  and  $\pi^0 n$  channels up to 2 GeV. The FA04 solution [29] takes into account the Erevan beam asymmetries up to 1.1 GeV [26] and the recent Bonn differential cross-sections [27] for the  $\pi^0 p$  reaction. The overall agreement is quite satisfactory especially for cross-sections. For the beam asymmetries (fig. 13), although the global trend is well reproduced, small discrepancies are observed for most energy bins.

MAID is a unitary isobar model designed to fit the pion photo- and electroproduction database. The model compared to our data is an extension of MAID2003 [30–32] which includes, for the  $\pi^0 p$  channel, the latest Bonn cross-sections and the present GRAAL beam asymmetries. As seen in the figures, the new MAID2005 solution is able to well reproduce the whole set of data and largely improves the previous solution (not shown in the figures). In particular, all observed structures are now qualitatively described. An important impact of the new data was to impose an empirical correction on the high-energy tail for  $\Delta(1232)P_{33}$  at  $W > 1600$  MeV [31]. Further studies are being pursued to better understand the origin of this unexpected modification.

The Bonn coupled-channel PWA is a combined analysis of photoproduction experiments with  $\pi N$ ,  $\eta N$ ,  $K \Lambda$  and  $K \Sigma$  final states [34]. As compared to the two previous models, it takes into account a larger database, with all polarization observable data, including our  $\pi^0 p$  beam asymmetries. This approach is an important step towards a global analysis of the complete photoproduction database.

The agreement for both observables is excellent for all energies and angles except for the lowest-energy bins. More generally, a fair agreement with the whole database has been obtained with 14  $N^*$  and 7  $\Delta^*$  resonances whose masses, widths and electromagnetic amplitudes are compatible with the PDG compilation. One of the main outcome of this model is the necessity to introduce several new resonances above 1800 MeV.

**Table 1.** Beam asymmetries for photon energies  $551 \text{ MeV} \leq E_\gamma \leq 1023 \text{ MeV}$ .

$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 551 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 578 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 615 \text{ MeV}$
		59.3	$0.417 \pm 0.149$	58.2	$0.330 \pm 0.068$
67.4	$0.431 \pm 0.050$	66.8	$0.463 \pm 0.034$	66.9	$0.546 \pm 0.036$
78.7	$0.592 \pm 0.041$	78.8	$0.601 \pm 0.033$	78.8	$0.665 \pm 0.033$
89.9	$0.670 \pm 0.040$	89.9	$0.721 \pm 0.032$	89.9	$0.774 \pm 0.033$
101.2	$0.694 \pm 0.044$	101.2	$0.783 \pm 0.031$	101.0	$0.807 \pm 0.033$
112.9	$0.736 \pm 0.041$	112.8	$0.750 \pm 0.035$	112.8	$0.835 \pm 0.035$
125.8	$0.836 \pm 0.049$	125.9	$0.810 \pm 0.038$	125.9	$0.830 \pm 0.037$
139.1	$0.796 \pm 0.061$	139.2	$0.800 \pm 0.047$	139.2	$0.743 \pm 0.046$
152.9	$0.698 \pm 0.119$	153.1	$0.696 \pm 0.092$	153.6	$0.543 \pm 0.079$

$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 666 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 703 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 736 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 769 \text{ MeV}$
44.7	$0.634 \pm 0.047$	44.7	$0.761 \pm 0.033$	44.7	$0.769 \pm 0.031$	44.7	$0.776 \pm 0.028$
52.8	$0.661 \pm 0.024$	52.8	$0.820 \pm 0.027$	52.7	$0.842 \pm 0.024$	52.7	$0.796 \pm 0.024$
61.7	$0.743 \pm 0.026$	61.8	$0.825 \pm 0.023$	61.7	$0.863 \pm 0.023$	61.7	$0.825 \pm 0.021$
70.4	$0.801 \pm 0.023$	70.4	$0.872 \pm 0.026$	70.4	$0.873 \pm 0.023$	70.3	$0.840 \pm 0.021$
78.6	$0.821 \pm 0.022$	78.6	$0.872 \pm 0.023$	78.6	$0.865 \pm 0.022$	78.6	$0.831 \pm 0.023$
86.3	$0.824 \pm 0.022$	86.3	$0.847 \pm 0.024$	86.3	$0.873 \pm 0.022$	86.2	$0.838 \pm 0.023$
94.0	$0.837 \pm 0.022$	94.0	$0.854 \pm 0.024$	94.0	$0.866 \pm 0.025$	94.0	$0.795 \pm 0.022$
102.2	$0.844 \pm 0.025$	102.2	$0.857 \pm 0.025$	102.2	$0.839 \pm 0.023$	102.2	$0.774 \pm 0.022$
110.8	$0.854 \pm 0.027$	110.8	$0.822 \pm 0.024$	110.8	$0.814 \pm 0.026$	110.8	$0.733 \pm 0.020$
119.3	$0.811 \pm 0.025$	119.2	$0.766 \pm 0.027$	119.2	$0.773 \pm 0.023$	119.2	$0.680 \pm 0.023$
130.7	$0.748 \pm 0.029$	130.7	$0.682 \pm 0.023$	130.6	$0.639 \pm 0.020$	130.7	$0.583 \pm 0.020$
139.7	$0.622 \pm 0.027$	139.8	$0.579 \pm 0.021$	139.9	$0.515 \pm 0.018$	139.9	$0.473 \pm 0.018$
150.2	$0.447 \pm 0.030$	150.2	$0.350 \pm 0.025$	150.3	$0.306 \pm 0.019$	150.4	$0.315 \pm 0.018$
161.0	$0.186 \pm 0.055$	161.0	$0.132 \pm 0.043$	160.9	$0.210 \pm 0.041$	160.9	$0.151 \pm 0.026$
170.7	$-0.019 \pm 0.133$	170.5	$0.115 \pm 0.157$	170.5	$0.041 \pm 0.083$	170.5	$0.138 \pm 0.084$

$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 801 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 826 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 862 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 895 \text{ MeV}$
44.7	$0.768 \pm 0.029$	44.8	$0.633 \pm 0.030$	44.8	$0.593 \pm 0.029$	44.8	$0.515 \pm 0.034$
52.7	$0.760 \pm 0.021$	52.8	$0.660 \pm 0.022$	52.8	$0.559 \pm 0.028$	52.7	$0.481 \pm 0.021$
61.7	$0.756 \pm 0.021$	61.7	$0.685 \pm 0.022$	61.8	$0.561 \pm 0.023$	61.7	$0.490 \pm 0.018$
70.4	$0.770 \pm 0.023$	70.4	$0.691 \pm 0.019$	70.4	$0.603 \pm 0.018$	70.4	$0.497 \pm 0.017$
78.6	$0.762 \pm 0.021$	78.6	$0.668 \pm 0.019$	78.6	$0.593 \pm 0.019$	78.6	$0.468 \pm 0.018$
86.3	$0.757 \pm 0.021$	86.3	$0.664 \pm 0.023$	86.3	$0.575 \pm 0.020$	86.3	$0.460 \pm 0.016$
94.0	$0.719 \pm 0.020$	94.0	$0.636 \pm 0.019$	94.0	$0.531 \pm 0.018$	94.0	$0.415 \pm 0.015$
102.2	$0.687 \pm 0.020$	102.2	$0.591 \pm 0.018$	102.2	$0.502 \pm 0.016$	102.3	$0.371 \pm 0.015$
110.8	$0.631 \pm 0.021$	110.8	$0.527 \pm 0.016$	110.8	$0.430 \pm 0.016$	110.9	$0.302 \pm 0.017$
119.2	$0.583 \pm 0.019$	119.2	$0.493 \pm 0.018$	119.2	$0.393 \pm 0.017$	119.3	$0.287 \pm 0.016$
131.0	$0.509 \pm 0.019$	131.0	$0.426 \pm 0.016$	130.9	$0.362 \pm 0.015$	130.9	$0.281 \pm 0.014$
139.9	$0.422 \pm 0.015$	140.0	$0.370 \pm 0.014$	140.0	$0.334 \pm 0.014$	140.0	$0.294 \pm 0.013$
150.3	$0.282 \pm 0.016$	150.4	$0.284 \pm 0.015$	150.4	$0.274 \pm 0.014$	150.4	$0.285 \pm 0.014$
160.8	$0.191 \pm 0.036$	160.8	$0.227 \pm 0.019$	160.8	$0.152 \pm 0.018$	160.7	$0.185 \pm 0.019$
170.4	$0.071 \pm 0.083$	170.4	$-0.052 \pm 0.060$	170.4	$0.020 \pm 0.059$	170.1	$0.046 \pm 0.081$

$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 929 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 961 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 992 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1023 \text{ MeV}$
44.7	$0.605 \pm 0.028$	44.8	$0.767 \pm 0.027$	44.8	$0.815 \pm 0.027$	44.9	$0.881 \pm 0.024$
52.6	$0.567 \pm 0.020$	52.5	$0.688 \pm 0.021$	52.5	$0.785 \pm 0.022$	52.5	$0.823 \pm 0.022$
61.7	$0.474 \pm 0.019$	61.5	$0.562 \pm 0.019$	61.5	$0.672 \pm 0.021$	61.5	$0.701 \pm 0.023$
70.4	$0.403 \pm 0.017$	70.3	$0.401 \pm 0.017$	70.3	$0.435 \pm 0.016$	70.2	$0.511 \pm 0.017$
78.6	$0.350 \pm 0.018$	78.6	$0.244 \pm 0.016$	78.6	$0.238 \pm 0.017$	78.5	$0.231 \pm 0.015$
86.3	$0.303 \pm 0.017$	86.4	$0.119 \pm 0.015$	86.3	$0.013 \pm 0.014$	86.3	$-0.076 \pm 0.015$
94.1	$0.223 \pm 0.018$	94.1	$-0.003 \pm 0.014$	94.1	$-0.169 \pm 0.014$	94.1	$-0.329 \pm 0.015$
102.3	$0.153 \pm 0.014$	102.3	$-0.079 \pm 0.014$	102.4	$-0.294 \pm 0.015$	102.4	$-0.460 \pm 0.017$
110.9	$0.110 \pm 0.013$	110.9	$-0.127 \pm 0.013$	111.0	$-0.305 \pm 0.014$	111.0	$-0.470 \pm 0.015$
119.4	$0.105 \pm 0.015$	119.5	$-0.081 \pm 0.014$	119.6	$-0.277 \pm 0.015$	119.7	$-0.388 \pm 0.016$
130.8	$0.168 \pm 0.015$	130.8	$0.025 \pm 0.012$	130.7	$-0.110 \pm 0.014$	130.7	$-0.237 \pm 0.013$
140.0	$0.229 \pm 0.012$	140.0	$0.131 \pm 0.011$	140.0	$0.034 \pm 0.011$	140.0	$-0.097 \pm 0.012$
150.4	$0.267 \pm 0.013$	150.3	$0.217 \pm 0.013$	150.3	$0.140 \pm 0.012$	150.3	$0.039 \pm 0.011$
160.7	$0.203 \pm 0.018$	160.6	$0.193 \pm 0.017$	160.6	$0.169 \pm 0.016$	160.6	$0.075 \pm 0.018$
170.1	$0.135 \pm 0.074$	170.2	$0.083 \pm 0.057$	170.0	$0.029 \pm 0.059$	169.8	$0.098 \pm 0.072$

**Table 2.** Beam asymmetries for photon energies  $1054 \text{ MeV} \leq E_\gamma \leq 1475 \text{ MeV}$ .

$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1054 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1085 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1114 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1144 \text{ MeV}$
44.9	$0.843 \pm 0.024$	45.0	$0.844 \pm 0.029$	45.0	$0.894 \pm 0.036$	45.0	$0.810 \pm 0.041$
52.5	$0.805 \pm 0.022$	52.5	$0.806 \pm 0.023$	52.7	$0.765 \pm 0.025$	52.7	$0.810 \pm 0.026$
61.5	$0.720 \pm 0.023$	61.5	$0.698 \pm 0.019$	61.5	$0.712 \pm 0.023$	61.5	$0.659 \pm 0.024$
70.2	$0.535 \pm 0.020$	70.2	$0.522 \pm 0.020$	70.1	$0.516 \pm 0.022$	70.1	$0.531 \pm 0.025$
78.5	$0.228 \pm 0.019$	78.4	$0.221 \pm 0.019$	78.4	$0.202 \pm 0.023$	78.4	$0.249 \pm 0.025$
86.3	$-0.143 \pm 0.016$	86.2	$-0.138 \pm 0.017$	86.2	$-0.192 \pm 0.024$	86.2	$-0.165 \pm 0.029$
94.1	$-0.433 \pm 0.017$	94.1	$-0.522 \pm 0.022$	94.0	$-0.523 \pm 0.025$	94.0	$-0.510 \pm 0.027$
102.4	$-0.575 \pm 0.018$	102.4	$-0.667 \pm 0.019$	102.3	$-0.684 \pm 0.031$	102.4	$-0.706 \pm 0.029$
111.0	$-0.581 \pm 0.018$	111.0	$-0.650 \pm 0.019$	110.9	$-0.667 \pm 0.023$	110.9	$-0.652 \pm 0.025$
119.9	$-0.486 \pm 0.020$	119.9	$-0.515 \pm 0.021$	119.6	$-0.552 \pm 0.027$	119.7	$-0.556 \pm 0.031$
130.6	$-0.310 \pm 0.015$	130.7	$-0.360 \pm 0.015$	130.9	$-0.362 \pm 0.019$	130.9	$-0.320 \pm 0.023$
140.0	$-0.181 \pm 0.013$	140.0	$-0.249 \pm 0.016$	140.0	$-0.260 \pm 0.016$	140.1	$-0.221 \pm 0.016$
150.4	$-0.068 \pm 0.013$	150.4	$-0.145 \pm 0.013$	150.4	$-0.178 \pm 0.016$	150.6	$-0.160 \pm 0.018$
160.7	$-0.004 \pm 0.015$	160.8	$-0.086 \pm 0.016$	160.9	$-0.095 \pm 0.025$	161.0	$-0.089 \pm 0.023$
169.9	$-0.003 \pm 0.054$	169.8	$-0.058 \pm 0.053$	169.8	$-0.117 \pm 0.079$	169.8	$-0.055 \pm 0.080$
$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1175 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1202 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1232 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1261 \text{ MeV}$
45.1	$0.886 \pm 0.042$	45.1	$0.832 \pm 0.045$	45.2	$0.839 \pm 0.048$	45.2	$0.878 \pm 0.049$
52.7	$0.783 \pm 0.031$	52.8	$0.797 \pm 0.030$	52.8	$0.881 \pm 0.032$	52.6	$0.858 \pm 0.031$
61.6	$0.700 \pm 0.026$	61.6	$0.723 \pm 0.027$	61.6	$0.725 \pm 0.029$	61.6	$0.721 \pm 0.031$
70.2	$0.544 \pm 0.027$	70.3	$0.572 \pm 0.027$	70.2	$0.532 \pm 0.027$	70.2	$0.576 \pm 0.028$
78.4	$0.299 \pm 0.033$	78.4	$0.356 \pm 0.027$	78.5	$0.446 \pm 0.027$	78.5	$0.390 \pm 0.028$
86.1	$-0.102 \pm 0.041$	86.2	$0.029 \pm 0.028$	86.2	$0.106 \pm 0.030$	86.2	$0.191 \pm 0.030$
94.0	$-0.406 \pm 0.028$	94.0	$-0.348 \pm 0.029$	94.0	$-0.197 \pm 0.041$	93.9	$-0.106 \pm 0.028$
102.3	$-0.693 \pm 0.028$	102.3	$-0.614 \pm 0.028$	102.3	$-0.480 \pm 0.041$	102.3	$-0.411 \pm 0.029$
110.9	$-0.649 \pm 0.027$	110.8	$-0.617 \pm 0.027$	110.9	$-0.515 \pm 0.027$	110.9	$-0.404 \pm 0.028$
119.9	$-0.477 \pm 0.036$	120.0	$-0.503 \pm 0.036$	120.4	$-0.383 \pm 0.033$	120.5	$-0.243 \pm 0.035$
130.9	$-0.252 \pm 0.022$	130.9	$-0.147 \pm 0.020$	130.9	$-0.032 \pm 0.018$	130.9	$0.045 \pm 0.018$
140.1	$-0.143 \pm 0.016$	140.1	$-0.024 \pm 0.016$	140.2	$0.063 \pm 0.016$	140.2	$0.181 \pm 0.015$
150.6	$-0.063 \pm 0.017$	150.6	$-0.005 \pm 0.018$	150.7	$0.093 \pm 0.017$	150.7	$0.172 \pm 0.016$
160.9	$-0.012 \pm 0.024$	161.0	$0.016 \pm 0.021$	161.0	$0.058 \pm 0.020$	161.0	$0.106 \pm 0.023$
169.8	$0.100 \pm 0.079$	169.7	$-0.064 \pm 0.097$	169.7	$-0.047 \pm 0.078$	169.7	$0.053 \pm 0.064$
$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1288 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1317 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1344 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1371 \text{ MeV}$
45.2	$0.965 \pm 0.051$	45.3	$0.949 \pm 0.056$	45.1	$0.990 \pm 0.069$	45.2	$0.846 \pm 0.035$
52.6	$0.886 \pm 0.037$	52.7	$0.865 \pm 0.035$	52.7	$0.864 \pm 0.039$	52.5	$0.846 \pm 0.035$
61.7	$0.761 \pm 0.030$	61.6	$0.790 \pm 0.032$	61.6	$0.752 \pm 0.044$	61.7	$0.722 \pm 0.035$
70.4	$0.619 \pm 0.028$	70.4	$0.622 \pm 0.030$	70.4	$0.593 \pm 0.034$	70.5	$0.568 \pm 0.032$
78.5	$0.481 \pm 0.028$	78.6	$0.477 \pm 0.029$	78.6	$0.502 \pm 0.030$	78.7	$0.495 \pm 0.029$
86.2	$0.257 \pm 0.031$	86.2	$0.282 \pm 0.029$	86.2	$0.338 \pm 0.028$	86.3	$0.440 \pm 0.034$
94.0	$-0.026 \pm 0.032$	94.0	$0.101 \pm 0.028$	94.0	$0.175 \pm 0.027$	94.0	$0.263 \pm 0.032$
102.2	$-0.277 \pm 0.028$	102.1	$-0.154 \pm 0.028$	102.2	$-0.013 \pm 0.027$	102.1	$0.025 \pm 0.027$
110.8	$-0.314 \pm 0.031$	110.8	$-0.190 \pm 0.033$	110.8	$-0.178 \pm 0.028$	110.6	$-0.060 \pm 0.029$
120.8	$-0.220 \pm 0.034$	120.8	$-0.024 \pm 0.041$	121.0	$0.017 \pm 0.035$	121.1	$0.115 \pm 0.035$
130.9	$0.144 \pm 0.018$	130.9	$0.218 \pm 0.018$	130.9	$0.316 \pm 0.018$	130.9	$0.386 \pm 0.019$
140.3	$0.267 \pm 0.018$	140.3	$0.304 \pm 0.016$	140.3	$0.394 \pm 0.016$	140.3	$0.444 \pm 0.018$
150.7	$0.253 \pm 0.017$	150.7	$0.281 \pm 0.018$	150.7	$0.333 \pm 0.020$	150.8	$0.382 \pm 0.018$
161.0	$0.124 \pm 0.019$	161.0	$0.178 \pm 0.022$	160.9	$0.187 \pm 0.018$	161.0	$0.211 \pm 0.019$
169.7	$-0.058 \pm 0.059$	169.7	$0.031 \pm 0.070$	169.8	$0.093 \pm 0.057$	169.7	$0.018 \pm 0.056$
$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1398 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1425 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1450 \text{ MeV}$	$\theta_{\text{cm}} (^{\circ})$	$E_\gamma = 1475 \text{ MeV}$
52.6	$0.894 \pm 0.046$	52.4	$0.918 \pm 0.039$	52.4	$0.744 \pm 0.048$	52.5	$0.829 \pm 0.051$
61.6	$0.722 \pm 0.039$	61.7	$0.633 \pm 0.038$	61.8	$0.518 \pm 0.046$	61.6	$0.521 \pm 0.075$
70.6	$0.615 \pm 0.032$	70.6	$0.578 \pm 0.034$	70.7	$0.514 \pm 0.038$	70.8	$0.457 \pm 0.045$
78.8	$0.498 \pm 0.030$	78.7	$0.475 \pm 0.034$	78.8	$0.552 \pm 0.034$	78.8	$0.484 \pm 0.036$
86.3	$0.409 \pm 0.029$	86.3	$0.460 \pm 0.031$	86.3	$0.481 \pm 0.031$	86.3	$0.508 \pm 0.033$
94.0	$0.284 \pm 0.026$	94.0	$0.405 \pm 0.032$	94.0	$0.321 \pm 0.029$	94.0	$0.469 \pm 0.031$
102.1	$0.155 \pm 0.029$	102.2	$0.243 \pm 0.026$	102.1	$0.279 \pm 0.031$	102.0	$0.364 \pm 0.033$
110.7	$0.028 \pm 0.031$	110.6	$0.106 \pm 0.029$	110.5	$0.157 \pm 0.034$	110.5	$0.244 \pm 0.037$
121.3	$0.138 \pm 0.035$	121.5	$0.179 \pm 0.044$	121.5	$0.308 \pm 0.038$	121.5	$0.231 \pm 0.044$
130.9	$0.418 \pm 0.021$	130.9	$0.418 \pm 0.019$	130.9	$0.501 \pm 0.025$	130.9	$0.520 \pm 0.023$
140.4	$0.473 \pm 0.020$	140.4	$0.516 \pm 0.020$	140.4	$0.578 \pm 0.020$	140.4	$0.606 \pm 0.021$
150.7	$0.423 \pm 0.019$	150.7	$0.435 \pm 0.020$	150.7	$0.461 \pm 0.021$	150.8	$0.505 \pm 0.022$
161.0	$0.241 \pm 0.020$	160.9	$0.262 \pm 0.021$	160.9	$0.273 \pm 0.023$	160.9	$0.323 \pm 0.025$
169.7	$0.072 \pm 0.053$	169.8	$0.128 \pm 0.063$	169.7	$0.057 \pm 0.055$	169.7	$0.023 \pm 0.077$

**Table 3.** Differential cross-sections  $d\sigma/d\Omega$  ( $\mu\text{b}/\text{sr}$ ) for photon energies  $555 \text{ MeV} \leq E_\gamma \leq 758 \text{ MeV}$ .

$\cos(\theta_{\text{cm}})$	$E_\gamma = 555 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 578 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 597 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 616 \text{ MeV}$
-0.97	$0.208 \pm 0.048$	-0.98	$0.294 \pm 0.057$	-0.98	$0.376 \pm 0.063$	-0.97	$0.505 \pm 0.073$
-0.90	$0.744 \pm 0.061$	-0.90	$0.634 \pm 0.059$	-0.90	$0.677 \pm 0.058$	-0.90	$0.710 \pm 0.057$
-0.80	$1.257 \pm 0.052$	-0.80	$1.190 \pm 0.053$	-0.80	$1.034 \pm 0.050$	-0.80	$1.000 \pm 0.048$
-0.69	$1.818 \pm 0.059$	-0.70	$1.617 \pm 0.067$	-0.70	$1.464 \pm 0.057$	-0.70	$1.534 \pm 0.054$
-0.59	$2.517 \pm 0.073$	-0.59	$2.113 \pm 0.071$	-0.60	$2.068 \pm 0.072$	-0.59	$1.926 \pm 0.069$
-0.49	$3.040 \pm 0.080$	-0.49	$2.562 \pm 0.083$	-0.49	$2.530 \pm 0.077$	-0.49	$2.409 \pm 0.076$
-0.39	$3.500 \pm 0.075$	-0.39	$3.014 \pm 0.076$	-0.39	$2.795 \pm 0.074$	-0.39	$2.675 \pm 0.074$
-0.29	$3.778 \pm 0.078$	-0.29	$3.381 \pm 0.087$	-0.29	$3.086 \pm 0.082$	-0.29	$2.901 \pm 0.075$
-0.19	$4.029 \pm 0.083$	-0.19	$3.582 \pm 0.084$	-0.19	$3.300 \pm 0.080$	-0.19	$3.027 \pm 0.076$
-0.10	$4.130 \pm 0.077$	-0.10	$3.569 \pm 0.095$	-0.10	$3.289 \pm 0.083$	-0.10	$3.308 \pm 0.096$
0.00	$4.179 \pm 0.088$	0.00	$3.601 \pm 0.095$	0.00	$3.207 \pm 0.087$	0.00	$3.122 \pm 0.087$
0.10	$4.063 \pm 0.082$	0.10	$3.498 \pm 0.091$	0.10	$3.152 \pm 0.081$	0.10	$2.979 \pm 0.080$
0.20	$3.815 \pm 0.075$	0.20	$3.206 \pm 0.085$	0.20	$2.916 \pm 0.075$	0.20	$2.712 \pm 0.074$
0.30	$3.576 \pm 0.070$	0.30	$3.015 \pm 0.080$	0.30	$2.808 \pm 0.075$	0.30	$2.647 \pm 0.076$
0.40	$3.215 \pm 0.066$	0.40	$2.718 \pm 0.078$	0.40	$2.335 \pm 0.069$	0.40	$2.226 \pm 0.070$
0.50	$2.797 \pm 0.104$	0.50	$2.319 \pm 0.082$	0.50	$2.197 \pm 0.077$	0.50	$1.925 \pm 0.063$
						0.60	$1.806 \pm 0.259$
$\cos(\theta_{\text{cm}})$	$E_\gamma = 634 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 652 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 670 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 688 \text{ MeV}$
-0.98	$0.529 \pm 0.076$	-0.98	$0.622 \pm 0.083$	-0.98	$0.793 \pm 0.090$	-0.97	$0.949 \pm 0.102$
-0.90	$0.797 \pm 0.061$	-0.90	$0.795 \pm 0.062$	-0.90	$0.991 \pm 0.072$	-0.90	$1.113 \pm 0.082$
-0.80	$1.021 \pm 0.047$	-0.80	$1.103 \pm 0.046$	-0.80	$1.195 \pm 0.044$	-0.80	$1.285 \pm 0.051$
-0.70	$1.471 \pm 0.056$	-0.69	$1.447 \pm 0.052$	-0.69	$1.497 \pm 0.054$	-0.70	$1.623 \pm 0.058$
-0.60	$1.887 \pm 0.066$	-0.60	$1.797 \pm 0.072$	-0.60	$1.969 \pm 0.071$	-0.60	$2.077 \pm 0.074$
-0.49	$2.264 \pm 0.081$	-0.49	$2.338 \pm 0.085$	-0.49	$2.417 \pm 0.076$	-0.49	$2.483 \pm 0.085$
-0.39	$2.661 \pm 0.077$	-0.39	$2.640 \pm 0.074$	-0.39	$2.654 \pm 0.080$	-0.39	$2.821 \pm 0.091$
-0.29	$2.883 \pm 0.078$	-0.29	$2.794 \pm 0.084$	-0.29	$3.073 \pm 0.084$	-0.29	$3.107 \pm 0.103$
-0.19	$3.076 \pm 0.079$	-0.19	$3.194 \pm 0.085$	-0.19	$3.101 \pm 0.092$	-0.19	$3.337 \pm 0.098$
-0.10	$3.164 \pm 0.096$	-0.09	$3.030 \pm 0.090$	-0.10	$3.051 \pm 0.093$	-0.10	$3.497 \pm 0.106$
0.00	$2.902 \pm 0.081$	0.00	$3.113 \pm 0.095$	0.00	$3.163 \pm 0.083$	0.00	$3.371 \pm 0.095$
0.10	$2.940 \pm 0.083$	0.10	$2.800 \pm 0.084$	0.10	$2.942 \pm 0.080$	0.10	$3.329 \pm 0.101$
0.20	$2.582 \pm 0.077$	0.20	$2.683 \pm 0.075$	0.20	$2.838 \pm 0.080$	0.20	$3.054 \pm 0.095$
0.30	$2.420 \pm 0.073$	0.30	$2.512 \pm 0.075$	0.30	$2.707 \pm 0.076$	0.30	$2.874 \pm 0.080$
0.40	$2.150 \pm 0.067$	0.40	$2.115 \pm 0.066$	0.40	$2.308 \pm 0.077$	0.40	$2.712 \pm 0.084$
0.50	$1.890 \pm 0.063$	0.50	$1.902 \pm 0.065$	0.50	$1.990 \pm 0.069$	0.50	$2.237 \pm 0.084$
0.60	$1.654 \pm 0.119$	0.60	$1.544 \pm 0.087$	0.60	$1.758 \pm 0.084$	0.60	$1.873 \pm 0.083$
$\cos(\theta_{\text{cm}})$	$E_\gamma = 705 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 723 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 740 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 758 \text{ MeV}$
-0.98	$1.228 \pm 0.114$	-0.98	$1.431 \pm 0.118$	-0.98	$1.642 \pm 0.128$	-0.98	$1.734 \pm 0.122$
-0.90	$1.412 \pm 0.099$	-0.90	$1.818 \pm 0.121$	-0.90	$2.067 \pm 0.133$	-0.90	$2.055 \pm 0.133$
-0.80	$1.694 \pm 0.059$	-0.80	$1.991 \pm 0.059$	-0.80	$2.196 \pm 0.061$	-0.80	$2.380 \pm 0.062$
-0.70	$2.023 \pm 0.063$	-0.70	$2.323 \pm 0.062$	-0.70	$2.477 \pm 0.061$	-0.70	$2.590 \pm 0.059$
-0.60	$2.400 \pm 0.078$	-0.60	$2.781 \pm 0.078$	-0.60	$2.789 \pm 0.080$	-0.60	$2.878 \pm 0.074$
-0.49	$2.808 \pm 0.091$	-0.49	$3.079 \pm 0.090$	-0.49	$3.373 \pm 0.100$	-0.49	$3.123 \pm 0.085$
-0.39	$3.333 \pm 0.097$	-0.39	$3.556 \pm 0.091$	-0.39	$3.694 \pm 0.089$	-0.39	$3.725 \pm 0.093$
-0.29	$3.565 \pm 0.101$	-0.29	$4.061 \pm 0.092$	-0.29	$4.031 \pm 0.094$	-0.29	$3.942 \pm 0.096$
-0.19	$3.958 \pm 0.102$	-0.19	$4.232 \pm 0.102$	-0.19	$4.428 \pm 0.100$	-0.19	$4.256 \pm 0.091$
-0.10	$4.067 \pm 0.104$	-0.10	$4.251 \pm 0.097$	-0.10	$4.547 \pm 0.100$	-0.10	$4.407 \pm 0.100$
0.00	$3.828 \pm 0.102$	0.00	$4.299 \pm 0.100$	0.00	$4.485 \pm 0.100$	0.00	$4.461 \pm 0.102$
0.10	$3.887 \pm 0.101$	0.10	$4.344 \pm 0.102$	0.10	$4.503 \pm 0.093$	0.10	$4.503 \pm 0.101$
0.20	$3.765 \pm 0.100$	0.20	$4.175 \pm 0.098$	0.20	$4.524 \pm 0.097$	0.20	$4.291 \pm 0.096$
0.30	$3.385 \pm 0.093$	0.30	$3.904 \pm 0.093$	0.30	$4.285 \pm 0.098$	0.30	$4.239 \pm 0.092$
0.40	$3.229 \pm 0.087$	0.40	$3.709 \pm 0.092$	0.40	$3.878 \pm 0.091$	0.40	$3.996 \pm 0.093$
0.50	$2.900 \pm 0.091$	0.50	$3.281 \pm 0.092$	0.50	$3.550 \pm 0.091$	0.50	$3.396 \pm 0.088$
0.60	$2.411 \pm 0.095$	0.60	$2.726 \pm 0.091$	0.60	$2.893 \pm 0.097$	0.60	$3.007 \pm 0.095$
				0.70	$2.208 \pm 0.363$	0.70	$2.363 \pm 0.204$



**Table 4.** Differential cross-sections  $d\sigma/d\Omega$  ( $\mu\text{b}/\text{sr}$ ) for photon energies  $774 \text{ MeV} \leq E_\gamma \leq 957 \text{ MeV}$ .

$\cos(\theta_{\text{cm}})$	$E_\gamma = 774 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 792 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 808 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 826 \text{ MeV}$
-0.98	$1.763 \pm 0.138$	-0.98	$1.810 \pm 0.127$	-0.98	$1.892 \pm 0.144$	-0.98	$1.889 \pm 0.129$
-0.90	$2.174 \pm 0.142$	-0.90	$2.401 \pm 0.153$	-0.90	$2.295 \pm 0.143$	-0.90	$2.381 \pm 0.151$
-0.80	$2.353 \pm 0.063$	-0.80	$2.498 \pm 0.061$	-0.80	$2.475 \pm 0.063$	-0.80	$2.628 \pm 0.060$
-0.70	$2.534 \pm 0.062$	-0.70	$2.577 \pm 0.059$	-0.70	$2.678 \pm 0.056$	-0.70	$2.675 \pm 0.054$
-0.60	$2.835 \pm 0.079$	-0.60	$2.918 \pm 0.078$	-0.60	$2.835 \pm 0.068$	-0.60	$2.758 \pm 0.066$
-0.49	$3.347 \pm 0.094$	-0.50	$3.098 \pm 0.086$	-0.49	$2.835 \pm 0.088$	-0.50	$2.741 \pm 0.079$
-0.39	$3.519 \pm 0.096$	-0.39	$3.409 \pm 0.094$	-0.39	$3.165 \pm 0.086$	-0.39	$2.995 \pm 0.080$
-0.29	$3.877 \pm 0.104$	-0.30	$3.614 \pm 0.092$	-0.30	$3.416 \pm 0.099$	-0.29	$3.057 \pm 0.085$
-0.19	$4.071 \pm 0.101$	-0.20	$3.848 \pm 0.097$	-0.20	$3.620 \pm 0.101$	-0.20	$3.191 \pm 0.099$
-0.10	$4.216 \pm 0.110$	-0.10	$3.928 \pm 0.107$	-0.10	$3.427 \pm 0.102$	-0.10	$3.129 \pm 0.097$
0.00	$4.288 \pm 0.102$	0.00	$4.029 \pm 0.106$	0.00	$3.602 \pm 0.107$	0.00	$3.197 \pm 0.097$
0.10	$4.239 \pm 0.100$	0.10	$3.972 \pm 0.101$	0.10	$3.532 \pm 0.100$	0.10	$3.113 \pm 0.099$
0.20	$4.364 \pm 0.104$	0.20	$3.763 \pm 0.107$	0.20	$3.434 \pm 0.098$	0.20	$2.981 \pm 0.084$
0.30	$3.987 \pm 0.107$	0.30	$3.736 \pm 0.104$	0.30	$3.283 \pm 0.096$	0.30	$2.788 \pm 0.093$
0.40	$3.763 \pm 0.100$	0.40	$3.348 \pm 0.098$	0.40	$2.974 \pm 0.092$	0.40	$2.619 \pm 0.083$
0.50	$3.423 \pm 0.103$	0.50	$2.932 \pm 0.093$	0.50	$2.541 \pm 0.089$	0.50	$2.193 \pm 0.081$
0.60	$2.598 \pm 0.100$	0.60	$2.522 \pm 0.092$	0.60	$2.049 \pm 0.085$	0.60	$1.752 \pm 0.076$
0.70	$2.046 \pm 0.194$	0.70	$1.862 \pm 0.155$	0.70	$1.781 \pm 0.141$	0.70	$1.311 \pm 0.108$
$\cos(\theta_{\text{cm}})$	$E_\gamma = 843 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 859 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 876 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 893 \text{ MeV}$
-0.98	$2.128 \pm 0.140$	-0.98	$2.054 \pm 0.153$	-0.98	$1.952 \pm 0.147$	-0.98	$1.771 \pm 0.130$
-0.90	$2.465 \pm 0.154$	-0.90	$2.422 \pm 0.153$	-0.90	$2.496 \pm 0.157$	-0.90	$2.304 \pm 0.146$
-0.80	$2.583 \pm 0.062$	-0.80	$2.547 \pm 0.061$	-0.80	$2.514 \pm 0.061$	-0.80	$2.472 \pm 0.059$
-0.70	$2.616 \pm 0.054$	-0.70	$2.487 \pm 0.054$	-0.70	$2.438 \pm 0.051$	-0.70	$2.425 \pm 0.051$
-0.60	$2.511 \pm 0.065$	-0.60	$2.532 \pm 0.062$	-0.60	$2.419 \pm 0.062$	-0.60	$2.335 \pm 0.060$
-0.50	$2.615 \pm 0.075$	-0.50	$2.541 \pm 0.072$	-0.50	$2.403 \pm 0.076$	-0.50	$2.307 \pm 0.069$
-0.39	$2.762 \pm 0.075$	-0.39	$2.501 \pm 0.073$	-0.39	$2.302 \pm 0.066$	-0.39	$2.283 \pm 0.067$
-0.30	$2.730 \pm 0.077$	-0.30	$2.761 \pm 0.077$	-0.30	$2.461 \pm 0.076$	-0.30	$2.226 \pm 0.070$
-0.20	$2.961 \pm 0.085$	-0.20	$2.625 \pm 0.082$	-0.20	$2.345 \pm 0.078$	-0.20	$2.139 \pm 0.074$
-0.10	$2.718 \pm 0.087$	-0.10	$2.658 \pm 0.078$	-0.10	$2.288 \pm 0.077$	-0.10	$2.136 \pm 0.080$
0.00	$2.884 \pm 0.085$	0.00	$2.558 \pm 0.087$	0.00	$2.226 \pm 0.079$	0.00	$2.030 \pm 0.082$
0.10	$2.807 \pm 0.090$	0.10	$2.438 \pm 0.079$	0.10	$2.144 \pm 0.078$	0.10	$1.843 \pm 0.073$
0.20	$2.703 \pm 0.084$	0.20	$2.419 \pm 0.089$	0.20	$2.004 \pm 0.078$	0.20	$1.682 \pm 0.064$
0.30	$2.468 \pm 0.081$	0.30	$2.089 \pm 0.079$	0.30	$1.888 \pm 0.072$	0.30	$1.526 \pm 0.063$
0.40	$2.241 \pm 0.081$	0.40	$2.001 \pm 0.072$	0.40	$1.696 \pm 0.068$	0.40	$1.560 \pm 0.060$
0.50	$1.847 \pm 0.072$	0.50	$1.627 \pm 0.066$	0.50	$1.461 \pm 0.062$	0.50	$1.189 \pm 0.052$
0.60	$1.615 \pm 0.069$	0.60	$1.361 \pm 0.062$	0.60	$1.137 \pm 0.056$	0.60	$1.075 \pm 0.049$
0.70	$1.173 \pm 0.098$	0.71	$1.163 \pm 0.095$	0.71	$1.031 \pm 0.089$	0.71	$1.002 \pm 0.085$
$\cos(\theta_{\text{cm}})$	$E_\gamma = 909 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 925 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 941 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 957 \text{ MeV}$
-0.98	$1.741 \pm 0.137$	-0.98	$1.433 \pm 0.107$	-0.98	$1.471 \pm 0.117$	-0.98	$1.535 \pm 0.122$
-0.90	$2.363 \pm 0.148$	-0.90	$2.293 \pm 0.150$	-0.90	$2.379 \pm 0.152$	-0.90	$2.376 \pm 0.152$
-0.80	$2.440 \pm 0.063$	-0.80	$2.555 \pm 0.063$	-0.80	$2.658 \pm 0.064$	-0.80	$2.671 \pm 0.064$
-0.70	$2.448 \pm 0.051$	-0.70	$2.429 \pm 0.050$	-0.70	$2.610 \pm 0.054$	-0.70	$2.659 \pm 0.053$
-0.60	$2.324 \pm 0.059$	-0.60	$2.419 \pm 0.056$	-0.60	$2.393 \pm 0.057$	-0.60	$2.577 \pm 0.063$
-0.50	$2.298 \pm 0.074$	-0.50	$2.236 \pm 0.073$	-0.50	$2.273 \pm 0.069$	-0.50	$2.366 \pm 0.074$
-0.39	$2.214 \pm 0.065$	-0.39	$2.166 \pm 0.064$	-0.39	$2.224 \pm 0.070$	-0.39	$2.278 \pm 0.069$
-0.30	$2.126 \pm 0.063$	-0.30	$2.117 \pm 0.068$	-0.30	$2.004 \pm 0.069$	-0.30	$1.992 \pm 0.064$
-0.20	$2.021 \pm 0.070$	-0.20	$1.916 \pm 0.068$	-0.20	$1.870 \pm 0.063$	-0.20	$1.861 \pm 0.072$
-0.10	$1.939 \pm 0.064$	-0.10	$1.765 \pm 0.066$	-0.10	$1.849 \pm 0.069$	-0.10	$1.673 \pm 0.067$
0.00	$1.814 \pm 0.066$	0.00	$1.645 \pm 0.071$	0.00	$1.708 \pm 0.066$	0.00	$1.537 \pm 0.064$
0.10	$1.690 \pm 0.065$	0.10	$1.553 \pm 0.064$	0.10	$1.540 \pm 0.070$	0.10	$1.504 \pm 0.063$
0.20	$1.619 \pm 0.062$	0.20	$1.475 \pm 0.064$	0.20	$1.413 \pm 0.060$	0.20	$1.333 \pm 0.059$
0.30	$1.401 \pm 0.061$	0.30	$1.332 \pm 0.056$	0.30	$1.312 \pm 0.054$	0.30	$1.354 \pm 0.058$
0.40	$1.393 \pm 0.060$	0.40	$1.160 \pm 0.052$	0.40	$1.237 \pm 0.050$	0.40	$1.226 \pm 0.054$
0.50	$1.157 \pm 0.050$	0.50	$1.144 \pm 0.051$	0.50	$1.175 \pm 0.051$	0.50	$1.344 \pm 0.057$
0.60	$1.091 \pm 0.053$	0.60	$1.148 \pm 0.053$	0.60	$1.285 \pm 0.057$	0.60	$1.374 \pm 0.065$
0.71	$0.996 \pm 0.088$	0.71	$0.943 \pm 0.092$	0.71	$1.160 \pm 0.106$	0.71	$1.578 \pm 0.125$

**Table 5.** Differential cross-sections  $d\sigma/d\Omega$  ( $\mu\text{b}/\text{sr}$ ) for photon energies  $973 \text{ MeV} \leq E_\gamma \leq 1096 \text{ MeV}$ .

$\cos(\theta_{\text{cm}})$	$E_\gamma = 973 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 989 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1005 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1021 \text{ MeV}$
-0.98	$1.602 \pm 0.124$	-0.98	$1.769 \pm 0.134$	-0.98	$1.753 \pm 0.122$	-0.98	$1.724 \pm 0.122$
-0.90	$2.456 \pm 0.158$	-0.90	$2.428 \pm 0.154$	-0.90	$2.473 \pm 0.161$	-0.90	$2.399 \pm 0.151$
-0.80	$2.787 \pm 0.067$	-0.80	$2.881 \pm 0.072$	-0.80	$3.063 \pm 0.074$	-0.80	$2.821 \pm 0.072$
-0.70	$2.739 \pm 0.053$	-0.70	$2.882 \pm 0.055$	-0.70	$3.063 \pm 0.053$	-0.70	$2.900 \pm 0.053$
-0.60	$2.733 \pm 0.063$	-0.60	$2.833 \pm 0.067$	-0.60	$2.846 \pm 0.063$	-0.60	$2.775 \pm 0.059$
-0.50	$2.428 \pm 0.073$	-0.50	$2.587 \pm 0.080$	-0.50	$2.721 \pm 0.086$	-0.50	$2.685 \pm 0.082$
-0.39	$2.385 \pm 0.076$	-0.39	$2.438 \pm 0.083$	-0.39	$2.435 \pm 0.082$	-0.39	$2.432 \pm 0.078$
-0.30	$2.092 \pm 0.069$	-0.30	$2.116 \pm 0.072$	-0.30	$2.165 \pm 0.068$	-0.29	$2.015 \pm 0.076$
-0.20	$1.921 \pm 0.071$	-0.20	$1.892 \pm 0.070$	-0.20	$1.950 \pm 0.075$	-0.20	$1.828 \pm 0.075$
-0.10	$1.651 \pm 0.066$	-0.10	$1.671 \pm 0.068$	-0.10	$1.640 \pm 0.071$	-0.10	$1.587 \pm 0.073$
0.00	$1.601 \pm 0.062$	0.00	$1.593 \pm 0.064$	0.00	$1.521 \pm 0.068$	0.00	$1.658 \pm 0.071$
0.10	$1.437 \pm 0.059$	0.10	$1.608 \pm 0.063$	0.10	$1.595 \pm 0.062$	0.10	$1.506 \pm 0.065$
0.20	$1.403 \pm 0.060$	0.20	$1.522 \pm 0.060$	0.20	$1.569 \pm 0.072$	0.20	$1.694 \pm 0.071$
0.30	$1.491 \pm 0.061$	0.30	$1.533 \pm 0.064$	0.30	$1.700 \pm 0.070$	0.30	$1.688 \pm 0.063$
0.40	$1.420 \pm 0.059$	0.40	$1.638 \pm 0.065$	0.40	$1.870 \pm 0.072$	0.40	$1.945 \pm 0.075$
0.50	$1.541 \pm 0.060$	0.50	$1.831 \pm 0.065$	0.50	$1.959 \pm 0.070$	0.50	$2.341 \pm 0.078$
0.60	$1.718 \pm 0.069$	0.61	$1.830 \pm 0.073$	0.61	$2.060 \pm 0.077$	0.61	$2.283 \pm 0.088$
0.71	$1.703 \pm 0.139$	0.71	$1.905 \pm 0.138$	0.71	$1.970 \pm 0.161$	0.71	$2.581 \pm 0.204$
$\cos(\theta_{\text{cm}})$	$E_\gamma = 1036 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1052 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1067 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1082 \text{ MeV}$
-0.98	$1.708 \pm 0.118$	-0.98	$1.662 \pm 0.112$	-0.98	$1.579 \pm 0.108$	-0.98	$1.658 \pm 0.107$
-0.90	$2.355 \pm 0.150$	-0.90	$2.350 \pm 0.146$	-0.90	$2.010 \pm 0.127$	-0.90	$1.829 \pm 0.114$
-0.80	$2.766 \pm 0.069$	-0.80	$2.552 \pm 0.066$	-0.80	$2.316 \pm 0.064$	-0.80	$2.096 \pm 0.060$
-0.70	$2.831 \pm 0.055$	-0.70	$2.480 \pm 0.050$	-0.70	$2.416 \pm 0.048$	-0.70	$2.159 \pm 0.051$
-0.60	$2.633 \pm 0.060$	-0.60	$2.468 \pm 0.057$	-0.60	$2.297 \pm 0.059$	-0.60	$1.964 \pm 0.054$
-0.50	$2.554 \pm 0.078$	-0.50	$2.293 \pm 0.083$	-0.50	$2.010 \pm 0.071$	-0.50	$1.834 \pm 0.062$
-0.39	$2.433 \pm 0.090$	-0.39	$2.273 \pm 0.094$	-0.39	$1.977 \pm 0.082$	-0.39	$1.692 \pm 0.076$
-0.29	$1.905 \pm 0.072$	-0.29	$1.834 \pm 0.069$	-0.29	$1.619 \pm 0.062$	-0.29	$1.500 \pm 0.063$
-0.20	$1.681 \pm 0.068$	-0.20	$1.559 \pm 0.064$	-0.20	$1.475 \pm 0.073$	-0.20	$1.298 \pm 0.053$
-0.10	$1.489 \pm 0.063$	-0.10	$1.441 \pm 0.063$	-0.10	$1.327 \pm 0.054$	-0.09	$1.257 \pm 0.057$
0.00	$1.433 \pm 0.062$	0.00	$1.444 \pm 0.060$	0.00	$1.341 \pm 0.060$	0.00	$1.179 \pm 0.061$
0.10	$1.514 \pm 0.061$	0.10	$1.561 \pm 0.064$	0.10	$1.347 \pm 0.058$	0.10	$1.381 \pm 0.071$
0.20	$1.689 \pm 0.075$	0.20	$1.657 \pm 0.064$	0.20	$1.502 \pm 0.072$	0.20	$1.597 \pm 0.066$
0.31	$1.779 \pm 0.070$	0.31	$1.890 \pm 0.078$	0.30	$1.806 \pm 0.081$	0.30	$1.792 \pm 0.070$
0.40	$2.122 \pm 0.083$	0.40	$2.221 \pm 0.084$	0.40	$2.073 \pm 0.079$	0.40	$2.117 \pm 0.077$
0.50	$2.358 \pm 0.078$	0.50	$2.324 \pm 0.077$	0.50	$2.294 \pm 0.077$	0.50	$2.196 \pm 0.076$
0.61	$2.515 \pm 0.084$	0.61	$2.521 \pm 0.084$	0.61	$2.415 \pm 0.085$	0.61	$2.427 \pm 0.090$
$\cos(\theta_{\text{cm}})$	$E_\gamma = 1096 \text{ MeV}$						
-0.98	$1.534 \pm 0.137$						
-0.90	$1.776 \pm 0.116$						
-0.80	$1.909 \pm 0.069$						
-0.70	$1.889 \pm 0.060$						
-0.60	$1.838 \pm 0.062$						
-0.51	$1.748 \pm 0.083$						
-0.39	$1.687 \pm 0.089$						
-0.29	$1.345 \pm 0.074$						
-0.19	$1.253 \pm 0.068$						
-0.09	$1.078 \pm 0.073$						
0.00	$1.318 \pm 0.083$						
0.10	$1.275 \pm 0.068$						
0.20	$1.486 \pm 0.076$						
0.30	$1.797 \pm 0.110$						
0.40	$2.082 \pm 0.101$						
0.50	$2.253 \pm 0.098$						
0.61	$2.310 \pm 0.113$						

**Table 6.** Differential cross-sections  $d\sigma/d\Omega$  ( $\mu\text{b}/\text{sr}$ ) for photon energies  $1107 \text{ MeV} \leq E_\gamma \leq 1418 \text{ MeV}$ .

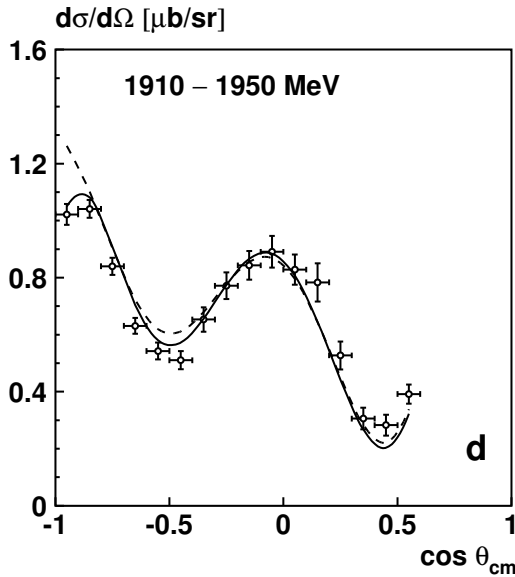
$\cos(\theta_{\text{cm}})$	$E_\gamma = 1107 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1137 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1167 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1196 \text{ MeV}$
-0.95	$1.382 \pm 0.070$	-0.95	$1.115 \pm 0.054$	-0.95	$1.141 \pm 0.056$	-0.95	$1.185 \pm 0.063$
-0.85	$1.439 \pm 0.062$	-0.85	$1.178 \pm 0.052$	-0.85	$1.072 \pm 0.049$	-0.85	$1.047 \pm 0.049$
-0.75	$1.571 \pm 0.078$	-0.75	$1.227 \pm 0.063$	-0.75	$1.028 \pm 0.054$	-0.75	$0.987 \pm 0.054$
-0.65	$1.472 \pm 0.117$	-0.65	$1.087 \pm 0.084$	-0.65	$0.873 \pm 0.079$	-0.65	$0.831 \pm 0.060$
-0.55	$1.400 \pm 0.103$	-0.55	$1.112 \pm 0.093$	-0.55	$0.935 \pm 0.092$	-0.55	$0.736 \pm 0.072$
-0.45	$1.276 \pm 0.097$	-0.45	$1.002 \pm 0.100$	-0.45	$0.811 \pm 0.071$	-0.45	$0.686 \pm 0.068$
-0.35	$1.208 \pm 0.085$	-0.35	$0.882 \pm 0.066$	-0.35	$0.777 \pm 0.065$	-0.35	$0.641 \pm 0.055$
-0.25	$1.107 \pm 0.082$	-0.25	$0.868 \pm 0.075$	-0.25	$0.740 \pm 0.059$	-0.25	$0.609 \pm 0.051$
-0.15	$1.041 \pm 0.065$	-0.15	$0.809 \pm 0.056$	-0.15	$0.705 \pm 0.065$	-0.15	$0.677 \pm 0.057$
-0.05	$1.004 \pm 0.074$	-0.05	$0.832 \pm 0.058$	-0.05	$0.794 \pm 0.058$	-0.05	$0.767 \pm 0.072$
0.05	$1.156 \pm 0.074$	0.05	$0.892 \pm 0.065$	0.05	$0.901 \pm 0.083$	0.05	$0.840 \pm 0.066$
0.15	$1.266 \pm 0.081$	0.15	$1.135 \pm 0.084$	0.15	$0.925 \pm 0.067$	0.15	$0.856 \pm 0.068$
0.25	$1.401 \pm 0.084$	0.25	$1.295 \pm 0.088$	0.25	$1.124 \pm 0.084$	0.25	$0.979 \pm 0.091$
0.35	$1.732 \pm 0.108$	0.35	$1.465 \pm 0.108$	0.35	$1.185 \pm 0.116$	0.35	$0.982 \pm 0.089$
0.45	$1.880 \pm 0.112$	0.45	$1.461 \pm 0.093$	0.45	$1.383 \pm 0.106$	0.45	$1.080 \pm 0.086$
0.55	$2.033 \pm 0.111$	0.55	$1.630 \pm 0.104$	0.55	$1.264 \pm 0.097$	0.55	$1.162 \pm 0.092$
0.65	$1.980 \pm 0.134$	0.65	$1.613 \pm 0.118$	0.65	$1.333 \pm 0.109$	0.65	$1.140 \pm 0.108$
$\cos(\theta_{\text{cm}})$	$E_\gamma = 1225 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1253 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1282 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1310 \text{ MeV}$
-0.95	$1.229 \pm 0.063$	-0.95	$1.337 \pm 0.071$	-0.95	$1.353 \pm 0.070$	-0.95	$1.446 \pm 0.072$
-0.85	$1.032 \pm 0.051$	-0.85	$1.169 \pm 0.054$	-0.85	$1.220 \pm 0.056$	-0.85	$1.278 \pm 0.064$
-0.75	$0.968 \pm 0.056$	-0.75	$0.987 \pm 0.054$	-0.75	$1.004 \pm 0.056$	-0.75	$1.038 \pm 0.059$
-0.65	$0.765 \pm 0.055$	-0.65	$0.821 \pm 0.060$	-0.65	$0.797 \pm 0.056$	-0.65	$0.831 \pm 0.067$
-0.55	$0.693 \pm 0.061$	-0.55	$0.700 \pm 0.063$	-0.55	$0.635 \pm 0.074$	-0.55	$0.651 \pm 0.056$
-0.45	$0.598 \pm 0.058$	-0.45	$0.588 \pm 0.061$	-0.45	$0.551 \pm 0.062$	-0.45	$0.633 \pm 0.072$
-0.35	$0.632 \pm 0.057$	-0.35	$0.594 \pm 0.061$	-0.35	$0.548 \pm 0.053$	-0.35	$0.542 \pm 0.050$
-0.25	$0.581 \pm 0.056$	-0.25	$0.574 \pm 0.051$	-0.25	$0.552 \pm 0.055$	-0.25	$0.591 \pm 0.051$
-0.15	$0.613 \pm 0.051$	-0.15	$0.604 \pm 0.052$	-0.15	$0.681 \pm 0.063$	-0.15	$0.738 \pm 0.062$
-0.05	$0.687 \pm 0.057$	-0.05	$0.633 \pm 0.054$	-0.05	$0.700 \pm 0.058$	-0.05	$0.702 \pm 0.059$
0.05	$0.733 \pm 0.064$	0.05	$0.714 \pm 0.068$	0.05	$0.656 \pm 0.053$	0.05	$0.677 \pm 0.065$
0.15	$0.775 \pm 0.073$	0.15	$0.755 \pm 0.074$	0.15	$0.817 \pm 0.073$	0.15	$0.757 \pm 0.071$
0.25	$0.892 \pm 0.079$	0.25	$0.817 \pm 0.077$	0.25	$0.812 \pm 0.079$	0.25	$0.724 \pm 0.071$
0.35	$0.911 \pm 0.083$	0.35	$0.865 \pm 0.083$	0.35	$0.752 \pm 0.081$	0.35	$0.671 \pm 0.071$
0.45	$0.944 \pm 0.088$	0.45	$0.734 \pm 0.072$	0.45	$0.778 \pm 0.096$	0.45	$0.650 \pm 0.076$
0.55	$0.933 \pm 0.092$	0.55	$0.844 \pm 0.087$	0.55	$0.751 \pm 0.072$	0.55	$0.677 \pm 0.065$
0.65	$1.004 \pm 0.102$	0.65	$0.962 \pm 0.111$	0.65	$0.931 \pm 0.119$	0.65	$0.916 \pm 0.129$
$\cos(\theta_{\text{cm}})$	$E_\gamma = 1338 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1364 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1392 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1418 \text{ MeV}$
-0.95	$1.480 \pm 0.071$	-0.95	$1.451 \pm 0.068$	-0.95	$1.513 \pm 0.071$	-0.95	$1.497 \pm 0.069$
-0.85	$1.330 \pm 0.058$	-0.85	$1.361 \pm 0.060$	-0.85	$1.376 \pm 0.068$	-0.85	$1.384 \pm 0.057$
-0.75	$1.109 \pm 0.060$	-0.75	$1.153 \pm 0.068$	-0.75	$1.069 \pm 0.055$	-0.75	$1.073 \pm 0.053$
-0.65	$0.794 \pm 0.054$	-0.65	$0.829 \pm 0.063$	-0.65	$0.794 \pm 0.058$	-0.65	$0.851 \pm 0.053$
-0.55	$0.694 \pm 0.060$	-0.55	$0.663 \pm 0.062$	-0.55	$0.639 \pm 0.053$	-0.55	$0.666 \pm 0.053$
-0.45	$0.617 \pm 0.082$	-0.45	$0.581 \pm 0.056$	-0.45	$0.621 \pm 0.061$	-0.45	$0.625 \pm 0.058$
-0.35	$0.562 \pm 0.051$	-0.35	$0.609 \pm 0.058$	-0.35	$0.606 \pm 0.059$	-0.35	$0.664 \pm 0.066$
-0.25	$0.695 \pm 0.065$	-0.25	$0.625 \pm 0.056$	-0.25	$0.668 \pm 0.053$	-0.25	$0.761 \pm 0.058$
-0.15	$0.721 \pm 0.060$	-0.15	$0.778 \pm 0.066$	-0.15	$0.823 \pm 0.071$	-0.15	$0.797 \pm 0.060$
-0.05	$0.708 \pm 0.056$	-0.05	$0.820 \pm 0.066$	-0.05	$0.853 \pm 0.075$	-0.05	$0.874 \pm 0.067$
0.05	$0.759 \pm 0.065$	0.05	$0.752 \pm 0.059$	0.05	$0.814 \pm 0.065$	0.05	$0.813 \pm 0.063$
0.15	$0.754 \pm 0.067$	0.15	$0.762 \pm 0.069$	0.15	$0.729 \pm 0.075$	0.15	$0.739 \pm 0.073$
0.25	$0.716 \pm 0.076$	0.25	$0.657 \pm 0.085$	0.25	$0.692 \pm 0.075$	0.25	$0.560 \pm 0.061$
0.35	$0.617 \pm 0.075$	0.35	$0.557 \pm 0.074$	0.35	$0.485 \pm 0.060$	0.35	$0.465 \pm 0.060$
0.45	$0.542 \pm 0.061$	0.45	$0.488 \pm 0.059$	0.45	$0.408 \pm 0.051$	0.45	$0.362 \pm 0.056$
0.55	$0.612 \pm 0.061$	0.55	$0.560 \pm 0.056$	0.55	$0.566 \pm 0.060$	0.55	$0.466 \pm 0.047$
0.65	$0.980 \pm 0.174$	0.65	$0.776 \pm 0.118$	0.65	$0.805 \pm 0.140$	0.65	$0.899 \pm 0.199$

**Table 7.** Differential cross-sections  $d\sigma/d\Omega$  ( $\mu\text{b}/\text{sr}$ ) for photon energies  $1444 \text{ MeV} \leq E_\gamma \leq 1541 \text{ MeV}$ .

$\cos(\theta_{\text{cm}})$	$E_\gamma = 1444 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1470 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1496 \text{ MeV}$	$\cos(\theta_{\text{cm}})$	$E_\gamma = 1521 \text{ MeV}$
-0.95	$1.331 \pm 0.062$	-0.95	$1.280 \pm 0.072$	-0.95	$1.104 \pm 0.062$	-0.95	$0.999 \pm 0.058$
-0.85	$1.321 \pm 0.057$	-0.85	$1.153 \pm 0.056$	-0.85	$1.149 \pm 0.056$	-0.85	$1.035 \pm 0.051$
-0.75	$1.016 \pm 0.050$	-0.75	$0.965 \pm 0.055$	-0.75	$0.879 \pm 0.050$	-0.75	$0.833 \pm 0.048$
-0.65	$0.772 \pm 0.047$	-0.65	$0.774 \pm 0.054$	-0.65	$0.690 \pm 0.049$	-0.65	$0.644 \pm 0.045$
-0.55	$0.598 \pm 0.048$	-0.55	$0.629 \pm 0.054$	-0.55	$0.554 \pm 0.048$	-0.55	$0.541 \pm 0.047$
-0.45	$0.567 \pm 0.051$	-0.45	$0.528 \pm 0.053$	-0.45	$0.533 \pm 0.054$	-0.45	$0.532 \pm 0.053$
-0.35	$0.606 \pm 0.052$	-0.35	$0.549 \pm 0.055$	-0.35	$0.657 \pm 0.068$	-0.35	$0.689 \pm 0.070$
-0.25	$0.730 \pm 0.057$	-0.25	$0.865 \pm 0.085$	-0.25	$0.734 \pm 0.069$	-0.25	$0.825 \pm 0.079$
-0.15	$0.925 \pm 0.075$	-0.15	$0.978 \pm 0.093$	-0.15	$0.891 \pm 0.083$	-0.15	$0.852 \pm 0.079$
-0.05	$0.826 \pm 0.080$	-0.05	$0.897 \pm 0.091$	-0.05	$0.904 \pm 0.087$	-0.05	$0.860 \pm 0.087$
0.05	$0.785 \pm 0.062$	0.05	$0.800 \pm 0.081$	0.05	$0.797 \pm 0.081$	0.05	$0.917 \pm 0.093$
0.15	$0.739 \pm 0.074$	0.15	$0.642 \pm 0.077$	0.15	$0.766 \pm 0.096$	0.15	$0.876 \pm 0.118$
0.25	$0.530 \pm 0.074$	0.25	$0.612 \pm 0.086$	0.25	$0.596 \pm 0.089$	0.25	$0.608 \pm 0.091$
0.35	$0.387 \pm 0.049$	0.35	$0.400 \pm 0.068$	0.35	$0.370 \pm 0.066$	0.35	$0.259 \pm 0.048$
0.45	$0.337 \pm 0.051$	0.45	$0.302 \pm 0.053$	0.45	$0.400 \pm 0.082$	0.45	$0.217 \pm 0.039$
0.55	$0.461 \pm 0.048$	0.55	$0.429 \pm 0.056$	0.55	$0.375 \pm 0.049$	0.55	$0.468 \pm 0.064$
0.65	$0.930 \pm 0.243$						

$\cos(\theta_{\text{cm}})$	$E_\gamma = 1541 \text{ MeV}$
-0.95	$0.963 \pm 0.071$
-0.85	$0.939 \pm 0.056$
-0.75	$0.808 \pm 0.057$
-0.65	$0.559 \pm 0.050$
-0.55	$0.533 \pm 0.058$
-0.45	$0.466 \pm 0.058$
-0.35	$0.613 \pm 0.084$
-0.25	$0.757 \pm 0.091$
-0.15	$0.787 \pm 0.098$
-0.05	$0.909 \pm 0.113$
0.05	$0.770 \pm 0.100$
0.15	$0.708 \pm 0.131$
0.25	$0.378 \pm 0.073$
0.35	$0.288 \pm 0.079$
0.45	$0.230 \pm 0.060$
0.55	$0.332 \pm 0.061$

**Fig. 18.** Differential cross-sections for  $W = 1910\text{--}1950 \text{ MeV}$  ( $E_\gamma = 1480\text{--}1560 \text{ MeV}$ ). The Bonn PWA fit with the new resonance  $N(2070)D_{15}$  is shown as solid line; the dashed line shows a fit when the  $5/2^-$  resonance is replaced by a  $7/2^-$  state.

Our results on both observables have brought essential constraints to the fit. On the one hand,  $\Sigma$  has allowed to restrict the number of contributing resonances. On the other hand,  $d\sigma/d\Omega$  helped to determine the spin of  $N(2070)D_{15}$ , one of the hypothetical new resonances. As shown in fig. 18, a spin-parity of  $5/2^-$  for the resonance is favored by the backward angular distribution measured at  $W = 1930 \text{ MeV}$  ( $E_\gamma = 1520 \text{ MeV}$ ).

Thanks to the recent improvements in the three models, an excellent overall agreement is now achieved despite the variety of observed structures *versus* energy, underlining the presence of a large number of resonant contributions. It should be noted that significant differences remain in the extracted resonance parameters, even for the most prominent ones. However, the procedures of extraction largely differ from one model to another, as well as the used database.

## 5 Summary

In this paper, we have presented new results on  $\pi^0$  photoproduction on the proton from 550 to 1500 MeV. Precise measurements for both differential cross-sections and

beam asymmetries have been obtained. These results complement the pion photoproduction database and extend the beam asymmetry measurements up to 1500 MeV. We have compared our results with three models, which are able to fit the whole set of data and bring interesting new information. Like for our previous results on  $\pi^+n$  and  $\eta p$ , the  $\pi^0 p$  beam asymmetry turned out to be a valuable information for the models.

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## References

1. J.P. Bocquet, V. Kuznetsov, D. Rebreyend (Editors), *Proceedings of the Workshop on the Physics of Excited Nucleons, NSTAR2004* (World Scientific, 2004).
2. *Review of Particle Physics 2004*, Phys. Lett. B **592**, 1 (2004).
3. R. Beck *et al.*, Phys. Rev. C **61**, 035204 (2000).
4. G. Blanpied *et al.*, Phys. Rev. C **64**, 025203 (2001).
5. J. Ajaka *et al.*, Phys. Rev. Lett. **81**, 1797 (1998).
6. J. Ajaka *et al.*, Phys. Lett. B **475**, 372 (2000).
7. O. Bartalini *et al.*, Phys. Lett. B **544**, 113 (2002).
8. Y. Assafiri *et al.*, Phys. Rev. Lett. **90**, 222001 (2003).
9. Z. Li, B. Saghai, Nucl. Phys. A **644**, 345 (1998).
10. N. Mukhopahyay, N. Mathur, Phys. Lett. B **444**, 7 (1998).
11. L. Tiator *et al.*, Phys. Rev. C **60**, 035210 (1999).
12. R. Workman *et al.*, Phys. Rev. C **62**, 048201 (2000).
13. V. Bellini *et al.*, Nucl. Instrum. Methods A **386**, 254 (1997).
14. F. Renard, Thesis, Université J. Fourier (Grenoble) 1999, unpublished.
15. C. Fry, MSci research project at Imperial College, LPSC internal report 2002, unpublished.
16. F.R. Arutyunian *et al.*, Teor. Fiz. **45**, 312 (1963).
17. L. Nicoletti, Thesis, Université J. Fourier (Grenoble) 2002, unpublished.
18. P. Levi Sandri *et al.*, Nucl. Instrum. Methods A **370**, 396 (1996).
19. M. Castoldi *et al.*, Nucl. Instrum. Methods A **403**, 22 (1998).
20. F. Ghio *et al.*, Nucl. Instrum. Methods A **404**, 71 (1998).
21. V. Kouznetsov *et al.*, Nucl. Instrum. Methods A **487**, 396 (2002).
22. D. Barancourt *et al.*, Nucl. Instrum. Methods A **388**, 226 (1997).
23. L. Mazzaschi *et al.*, Nucl. Instrum. Methods A **436**, 441 (1994).
24. F. Renard *et al.*, Phys. Lett. B **528**, 215 (2002).
25. The full data set is available at the website <http://lpsc.in2p3.fr/GRAAL1>.
26. F.V. Adamian *et al.*, Phys. Rev. C **63**, 054606 (2001).
27. O. Bartholomy *et al.*, Phys. Rev. Lett. **94**, 012003 (2005).
28. R.A. Arndt *et al.*, Phys. Rev. C **66**, 055213 (2002).
29. SAID-FA04, website <http://gwdac.phys.gwu.edu>.
30. L. Tiator *et al.*, Eur. Phys. J. A **19**, S01, 55 (2004).
31. S.S. Kamalov, L. Tiator, MAID2005, private communication.
32. L. Tiator *et al.*, *Proceedings of the Workshop on the Physics of Excited Nucleons, NSTAR2005*, to be published by World Scientific.
33. A.V. Anisovich *et al.*, Phys. Rev. C **68**, 045202 (2003).
34. A.V. Anisovich *et al.*, Eur. Phys. J. A **25**, 427 (2005).